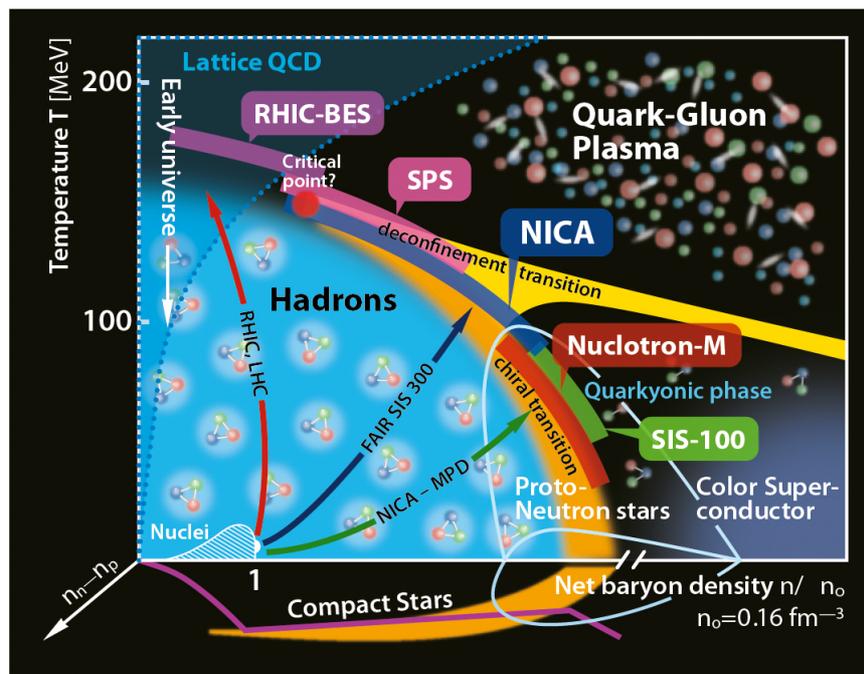


Nuclear Equation of State

Introduction

The nuclear equation of state (**EoS**) describes the **energy per nucleon** as a function of the neutron (ρ_n) and proton (ρ_p) densities of an **uniform and infinite system at zero temperature** that interact only via the residual strong interaction, or nuclear force. (*Neglect the Coulomb interaction*)

Ideal system : in the interior of nuclei and in cold neutron stars.



Nuclear energy density functional (**EDF**) constitutes a unique tool to reliably and consistently access bulk ground state and collective excited state properties of atomic nuclei along the nuclear chart as well as the **EoS around saturation density**.

Phenomenology

saturation density $\rho_0 = 0.16 \text{ fm}^{-3}$

Ground state

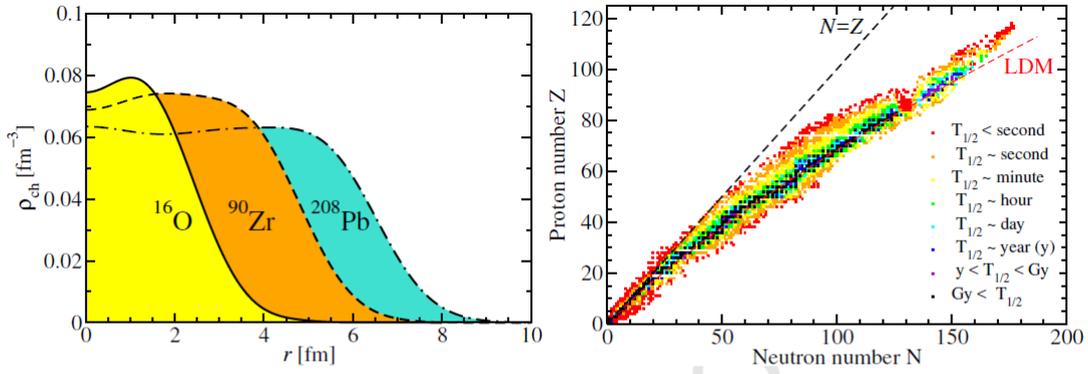


Figure 1: Left panel: experimental charge density (ρ_{ch}) as a function of the distance to the center of the nucleus (r , assuming spherical symmetry) derived from available data on elastic electron scattering for ^{16}O , ^{90}Zr and ^{208}Pb [16]. Right panel: Chart of nuclides classified by half-life. Data taken from NUBASE16 [17].

The density in the interior of very different nuclei is almost the same $\rho_{0p} = 0.07 \text{ fm}^{-3} \rightarrow$ a saturation mechanism (equilibrium) $\rightarrow \langle r_p^2 \rangle^{1/2} \approx 1.2Z^{1/3}$.

Stable nuclei are not far to be symmetric + assuming the nuclear force is isospin invariant $\rightarrow \rho_{0p} \approx \rho_{0n} \rightarrow \langle r^2 \rangle^{1/2} \approx 0.9A^{1/3}$

The most famous macroscopic model : **liquid drop model**
(neglecting pairing effects among others)

$$M(A, Z) = m_p Z + m_n (A - Z) - B(A, Z), \quad \text{where}$$

$$B(A, Z) = (a_V - a_S A^{-1/3})A - a_C \frac{Z(Z-1)}{A^{1/3}} - (a_A - a_{SA} A^{-1/3}) \frac{(A-2Z)^2}{A}$$

Symmetry energy term

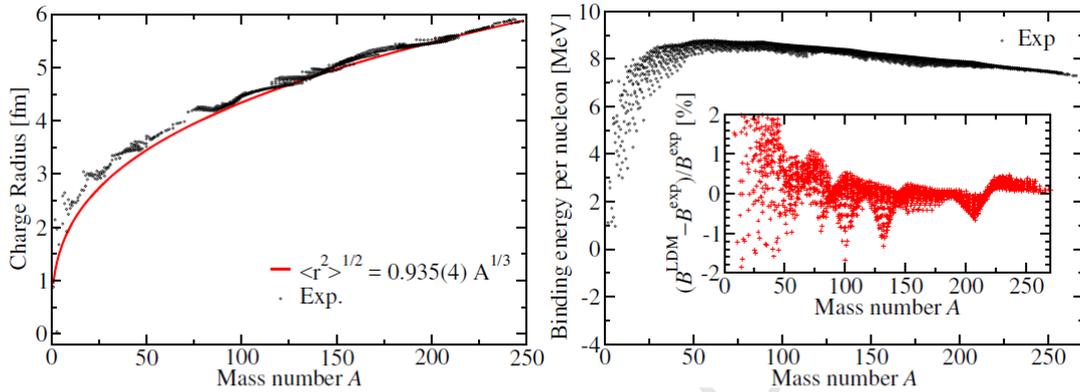


Figure 2: Left panel: root mean square charge radii of all measured nuclei as a function of the mass number A reported in Ref.[18]. Right panel: binding energy per particle of all measured nuclei by 2016 that were compiled in AME16 [21]. The inset shows the relative difference between the liquid drop model described in the text and the experimental data in %.

Liquid drop model \rightarrow diffuse surface ¹ \rightarrow droplet model (DM) \rightarrow prediction of neutron skin thickness ($\Delta r_{np} \equiv \langle r_n^2 \rangle^{1/2} - \langle r_p^2 \rangle^{1/2}$)

$$\Delta r_{np}^{\text{DM}} = \frac{2\langle r^2 \rangle^{1/2}}{3} \frac{a_{AS}}{a_A} (I - I_C) + \Delta r_{np}^C$$

Relative neutron excess : $I \equiv (N - Z)/A$; Coulomb correction I_C ; shift in the neutron skin due to the Coulomb interaction : Δr_{np}^C

As isospin asymmetries $\delta \equiv (\rho_n - \rho_p)/(\rho_n + \rho_p)$ are relatively small, an expansion for $\delta \rightarrow 0$:

$$e(\rho, \delta) = e(\rho, 0) + S(\rho)\delta^2 + \mathcal{O}[\delta^4], \quad \text{where}$$

$$\text{EoS of symmetric matter: } e(\rho, 0) = e(\rho_0, 0) + \frac{1}{2}K_0 \left(\frac{\rho - \rho_0}{3\rho_0} \right)^2 + \mathcal{O}[(\rho - \rho_0)^3],$$

$$\text{symmetry energy: } S(\rho) \equiv \left. \frac{\partial e(\rho, \delta)}{\partial \delta} \right|_{\delta=0} = J + L \frac{\rho - \rho_0}{3\rho_0} + \frac{1}{2}K_{\text{sym}} \left(\frac{\rho - \rho_0}{3\rho_0} \right)^2 + \mathcal{O}[(\rho - \rho_0)^3]$$

Incompressibility of symmetric nuclear matter K_0 ; symmetry energy at saturation

$J \equiv S(\rho_0)$; the slope of the symmetry energy

at saturation L ; the incompressibility (or curvature) of the symmetry energy at saturation

K_{sym}

Neutron matter EoS : $\delta = 1$, $S(\rho) \approx e(\rho, 1) - e(\rho, 0)$

References

1. W.D Myers and W.J Swiatecki. The nuclear droplet model for arbitrary shapes. *Annals of Physics*, 84(1):186 – 210, 1974. ↩