

# Predicting hypernuclei based on chiral interactions

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Seminar at Tongji University, Shanghai, China

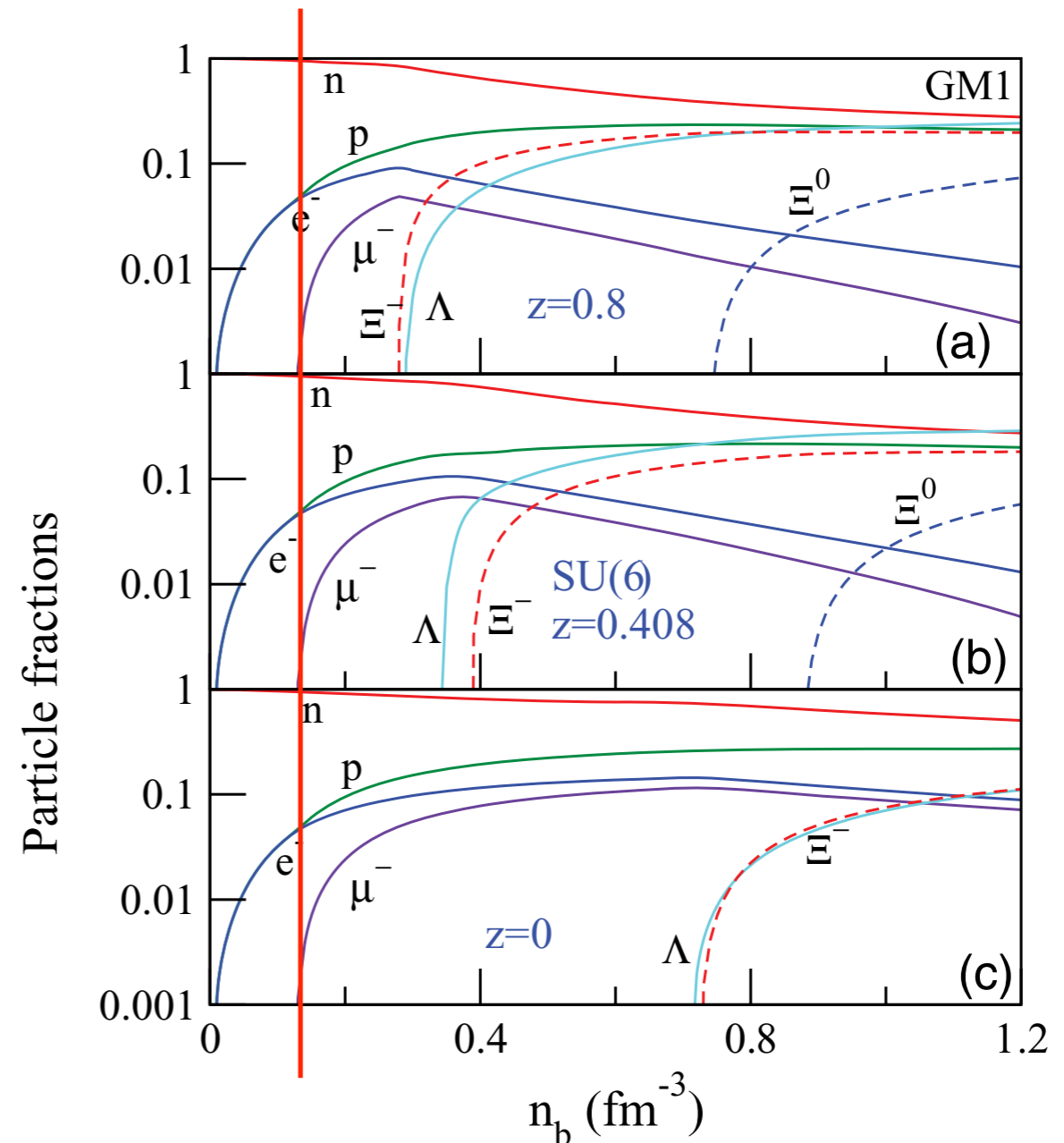
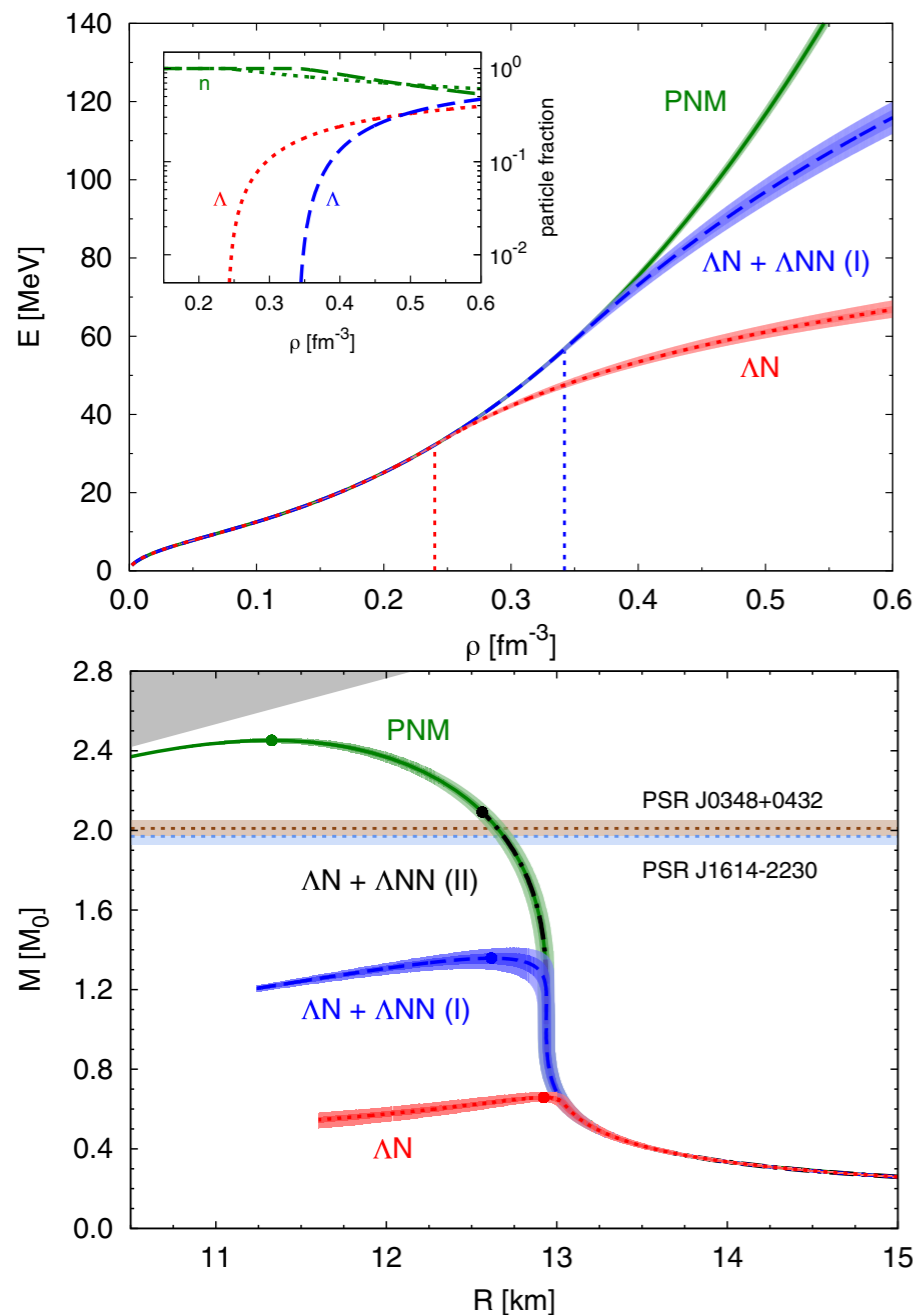
- Motivation
- YN & YY interactions
- J-NCSM & SRG evolution of (hyper-)nuclear interactions
- Determination of CSB contact interactions and  $\Lambda n$  scattering length  
& application to  $A = 7$  and 8 hypernuclei
- Uncertainty of  $\Lambda$  separation energies & chiral YNN interactions
- $S = -2$  hypernuclei: predictions for  $A \leq 6$
- SRG & long-range corrections of two-nucleon densities
- Conclusions & Outlook

in collaboration with Johann Haidenbauer, Hoai Le, Ulf Meißner, **Xiang-Xiang Sun**

# Hypernuclear interactions

## Why is understanding hypernuclear interactions interesting?

- *hyperon contribution to the EOS, neutron stars, supernovae*
- *"hyperon puzzle"*
- *$\Lambda$  as probe to nuclear structure*
- *flavor dependence of baryon-baryon interactions*



# Hypernuclei

Only few YN data. Hypernuclear data provides additional constraints.

- $\Lambda$ N interactions are generally weaker than the NN interaction
  - naively: **core nucleus + hyperons**
  - „separation energies“ are quite independent from NN(+3N) interaction
- no Pauli blocking of  $\Lambda$  in nuclei
  - good to study nuclear structure
  - even light hypernuclei exist in **several spin states**
- **non-trivial constraints** on the YN interaction even from lightest ones
- size of **YNN** interactions?  
need to include  **$\Lambda$ - $\Sigma$  conversion!**



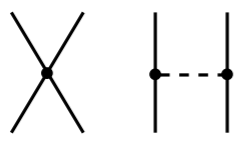


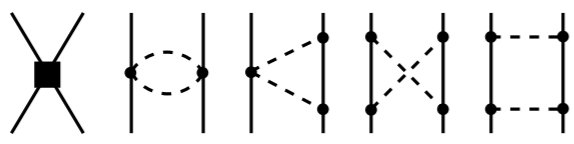


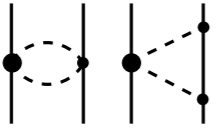
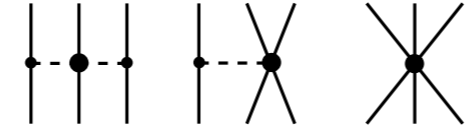

(from Panda@FAIR web page)

# Chiral NN & YN & YY interactions

## EFT based approaches

Chiral EFT implements **chiral symmetry of QCD**

- symmetries constrain exchanges of Goldstone bosons
- relations of two- and three- and more-baryon interactions
- breakdown scale  $\approx 600 - 700 \text{ MeV}$
- Semi-local momentum regularization (SMS) up to N<sup>2</sup>LO (for YN)

	BB force	3B force	4B force	
LO				<b>5(+1) NN/YN (YY)</b> short range parameters
NLO				<b>23(+5) NN/YN (YY)</b> short range parameters
N <sup>2</sup> LO				no additional contact terms in NN/YN

(adapted from Epelbaum, 2008)

Retain flexibility to adjust to data due to counter terms

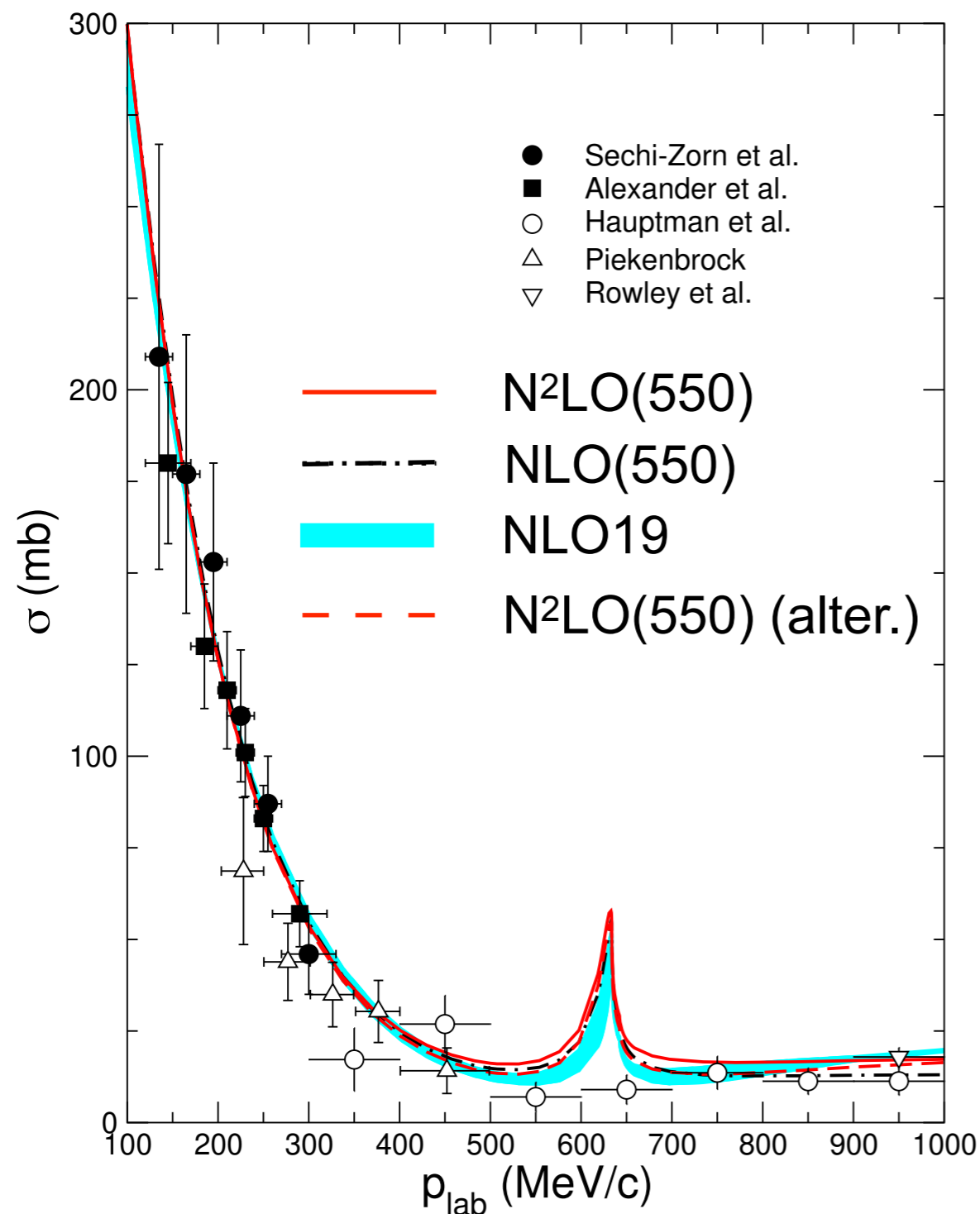
**Regulator required** — cutoff/different orders often used to estimate uncertainty

$\Lambda - \Sigma$  ( $\Lambda\Lambda - \Sigma\Sigma - \Xi N$ ) **conversion** is explicitly included (3BFs only in N<sup>2</sup>LO)

# SMS NLO/N<sup>2</sup>LO interaction

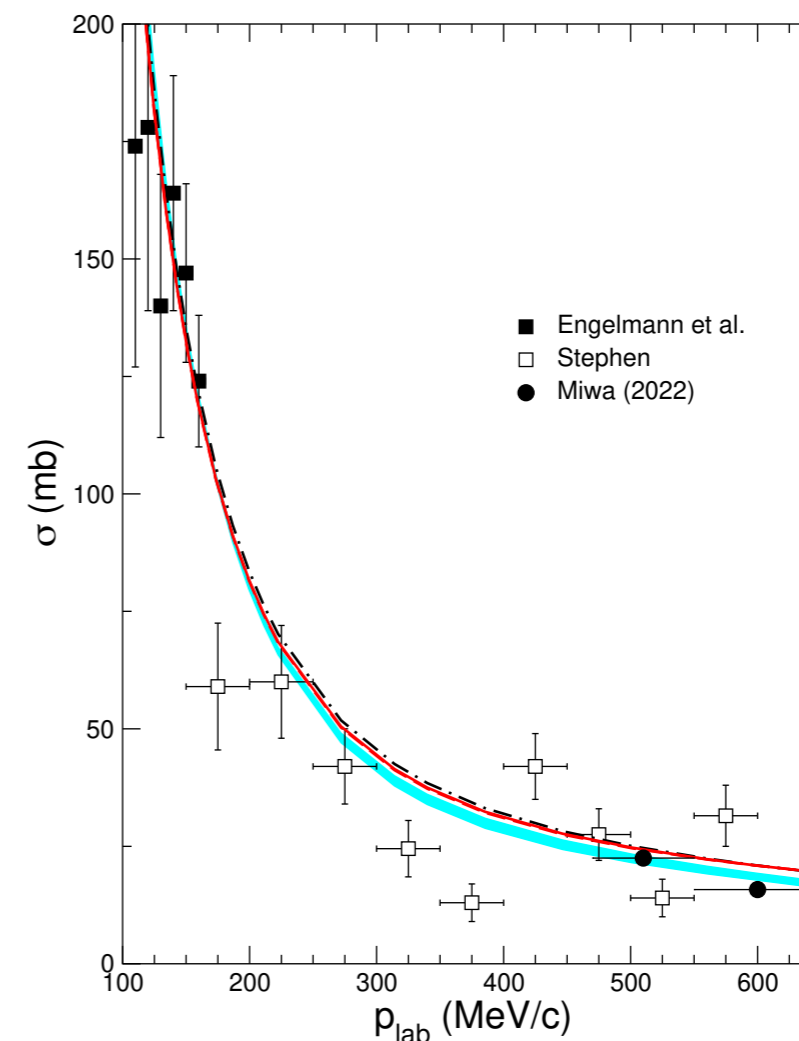
Selected results (show  $\Lambda = 550$  MeV, others are very similar in quality)

$\Lambda p \rightarrow \Lambda p$



- most relevant cross sections very similar in NLO and N<sup>2</sup>LO
- similar to NLO19
- alternative fit (see later)

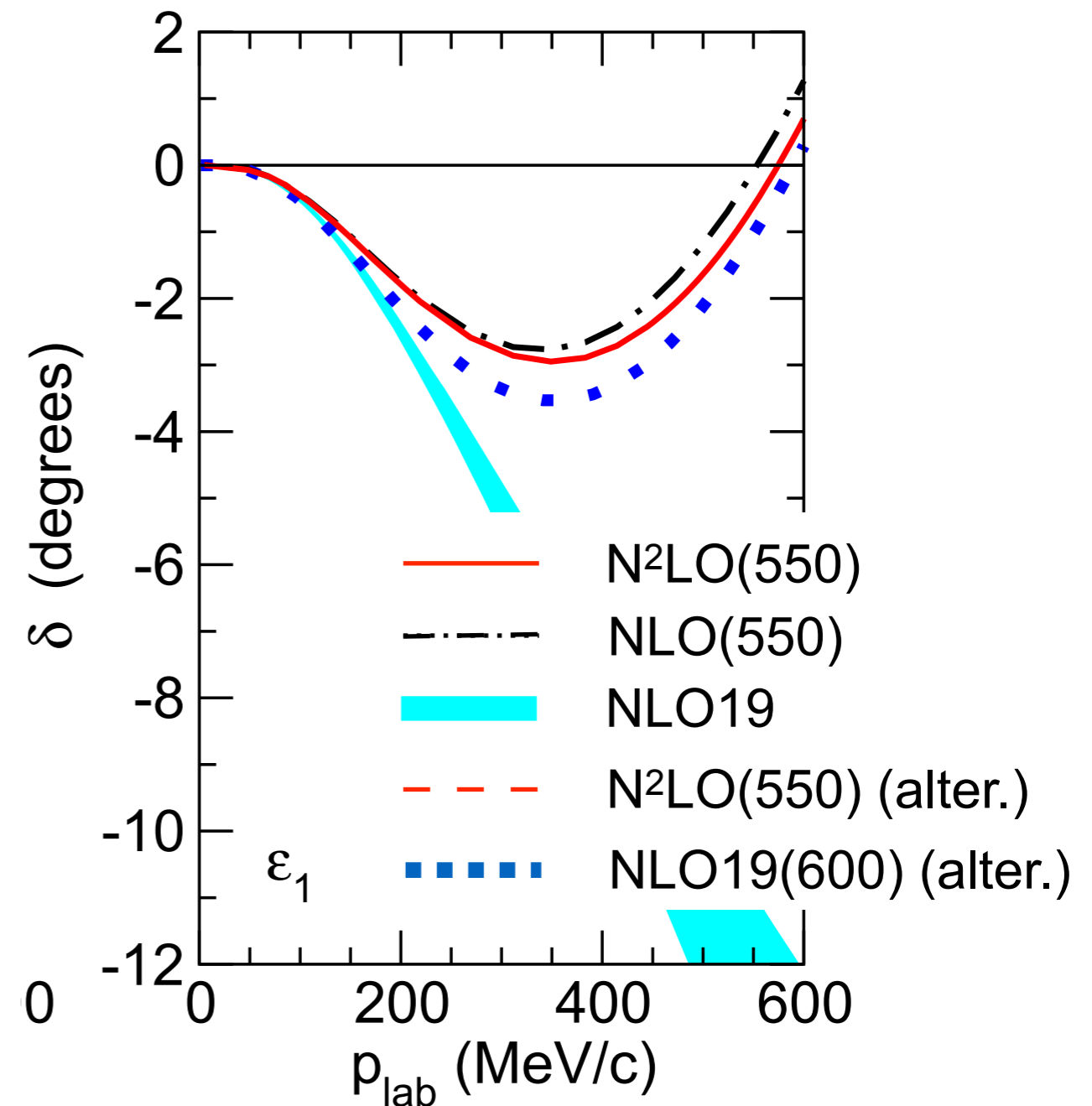
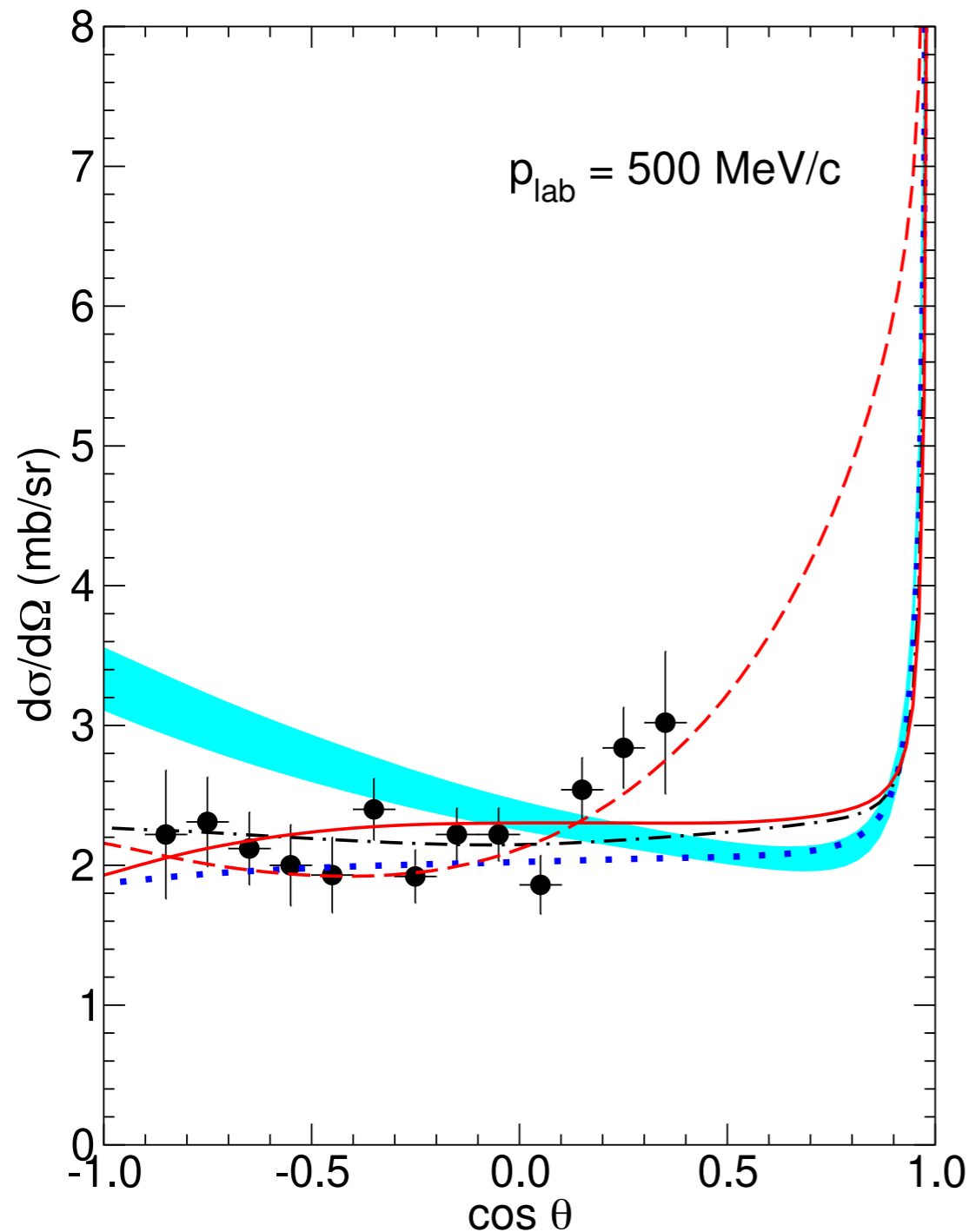
$\Sigma^- p \rightarrow \Lambda n$



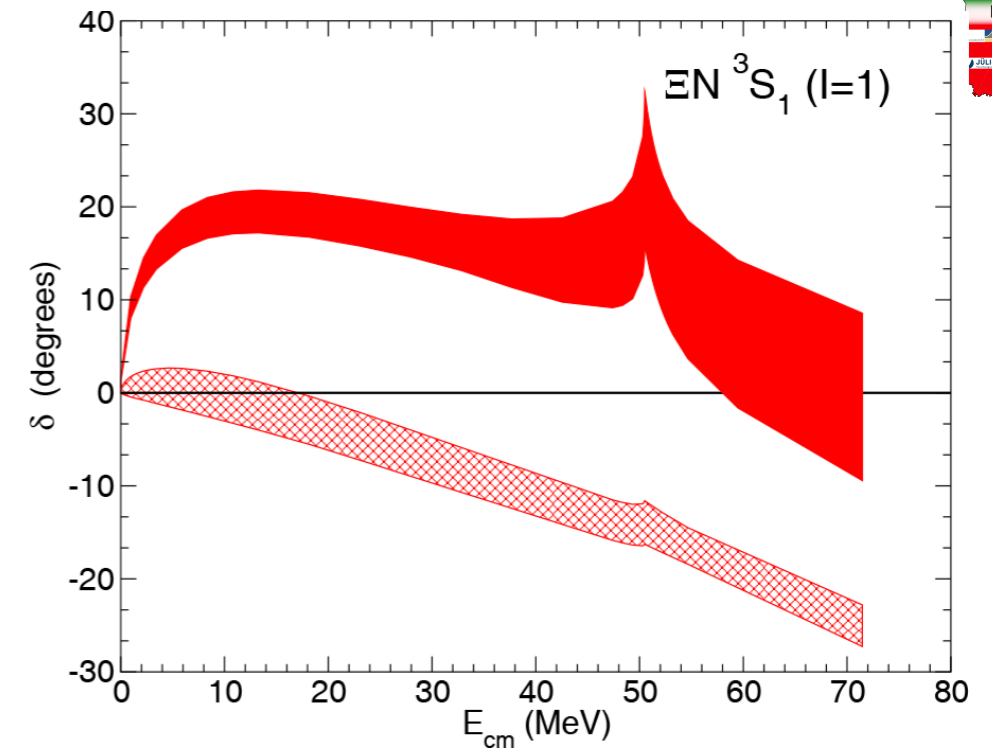
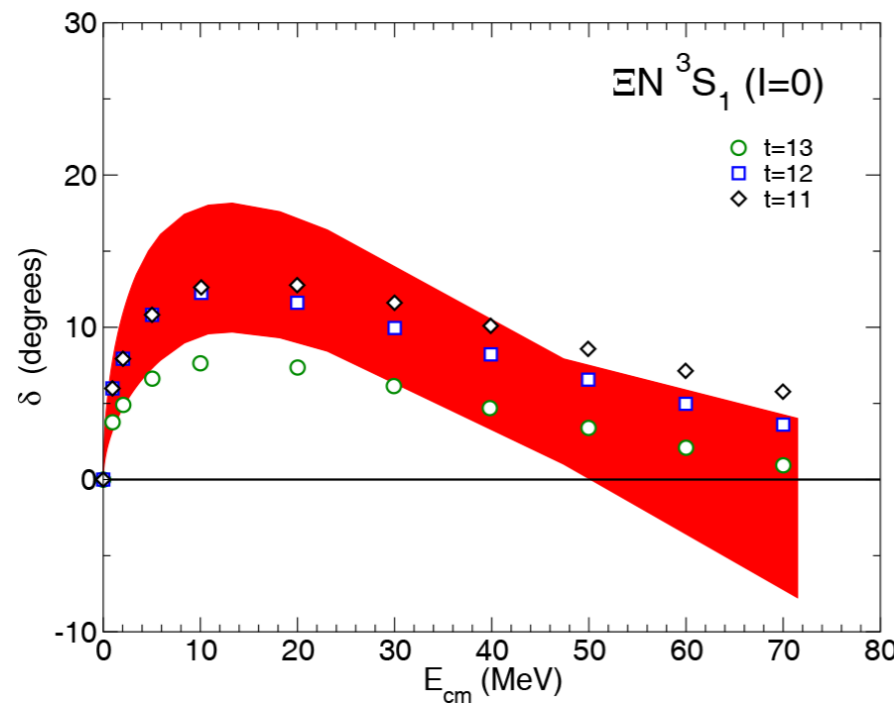
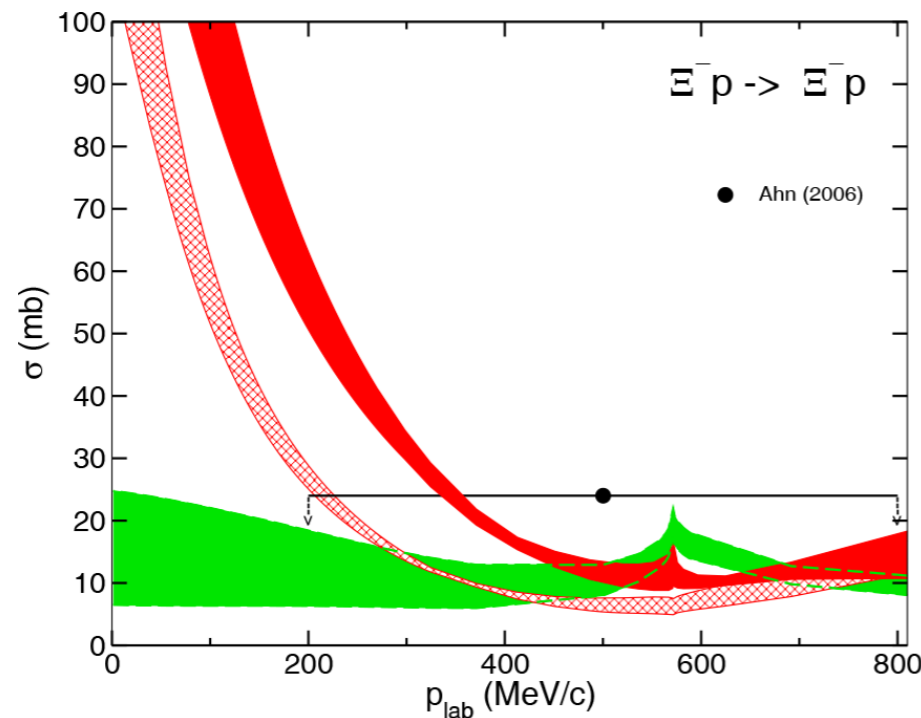
# SMS NLO/N<sup>2</sup>LO interaction


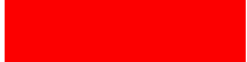

new data (Miwa(2022)) at higher energies provides new constraints!

$$\Sigma^+ p \rightarrow \Sigma^+ p$$



# YY interaction



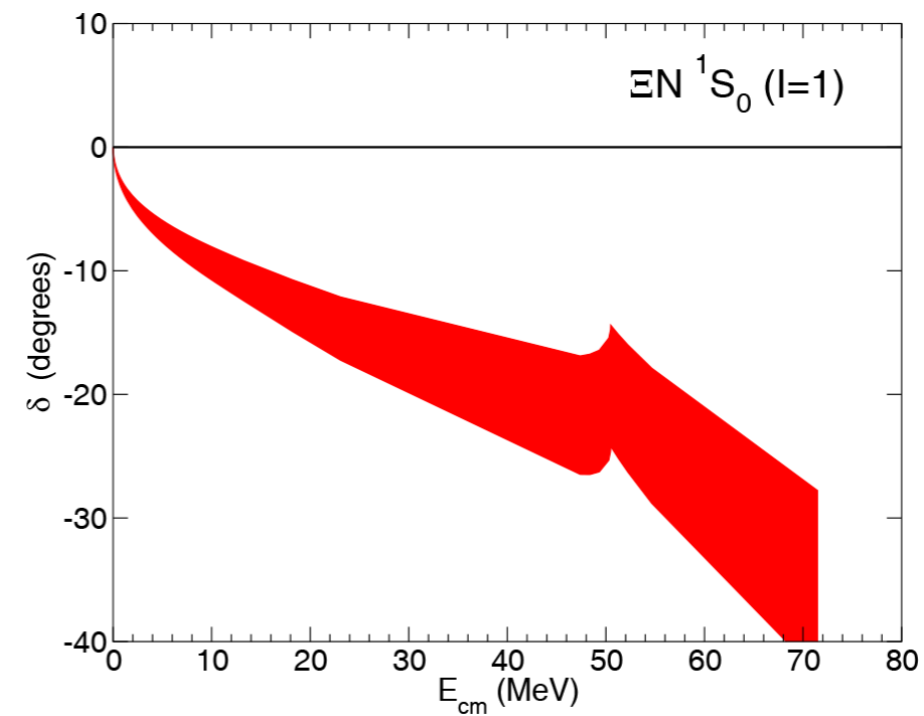
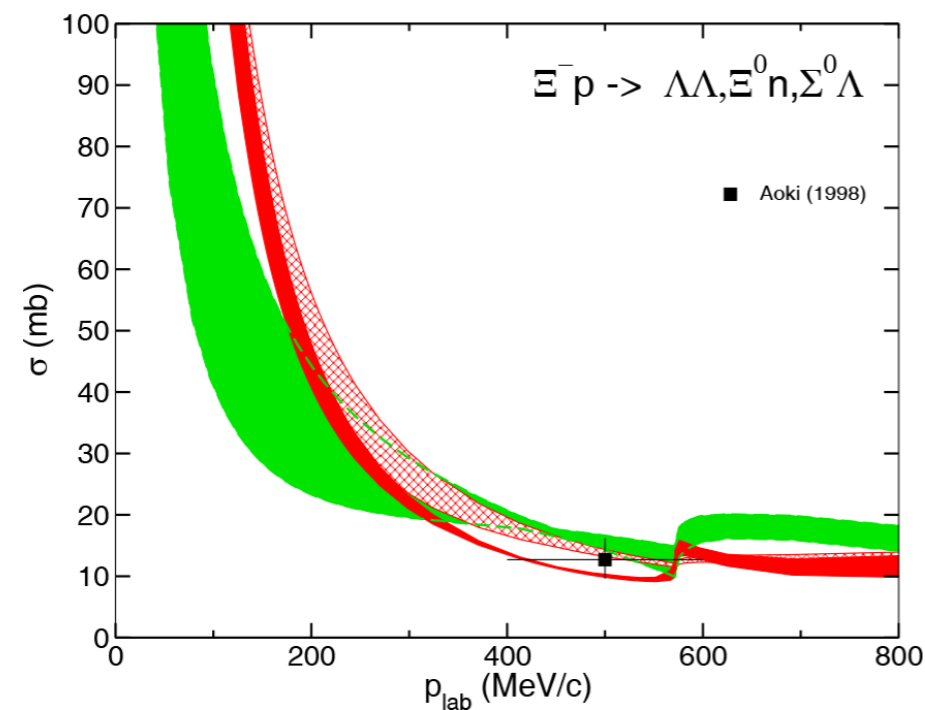
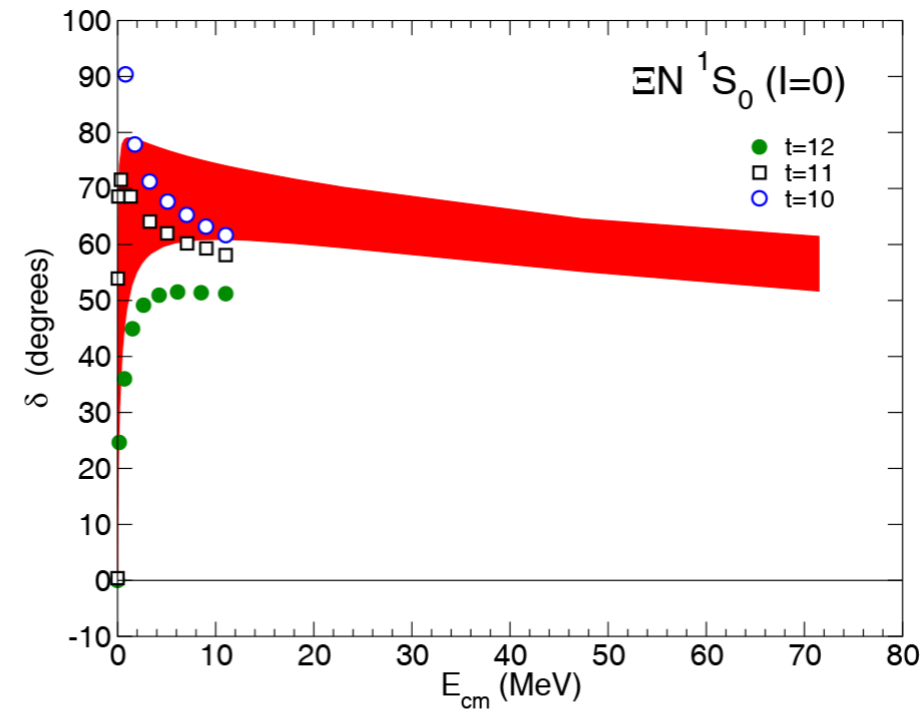
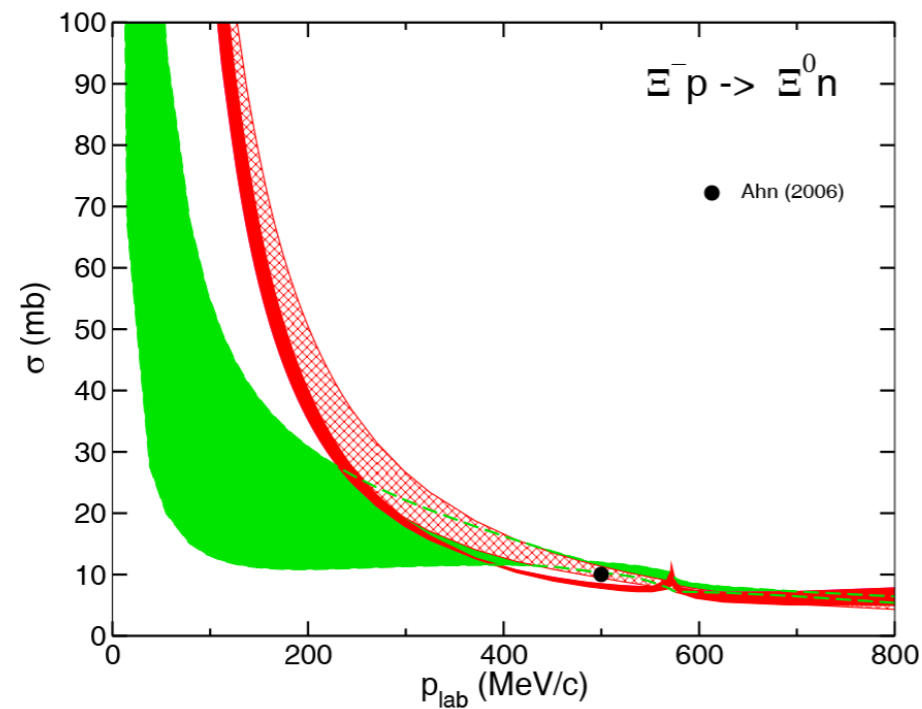
 LO  
 NLO  
 NLO(2016)  
 (Haidenbauer et al., 2019)



adjusted to data & LQCD (HAL QCD)

updated version consistent with  $\Xi$ -nuclei (only change in  $\Xi N \ ^3S_1$ )

# YY interaction



LO



NLO



NLO(2016)

(Haidenbauer et al., 2019)

Solve the Schrödinger equation using **HO states**

Two ingredients are necessary:

- **cfp** — antisymmetrized states for nucleons
- **transition coefficients** to separate off NN, YN, 3N and YNN

Schrödinger equation

$$\langle \text{blue circle with red dot} | H | \text{blue circle with red dot} \rangle \langle \text{blue circle with red dot} | \Psi \rangle = E \langle \text{blue circle with red dot} | \Psi \rangle$$

e.g. for YN interaction

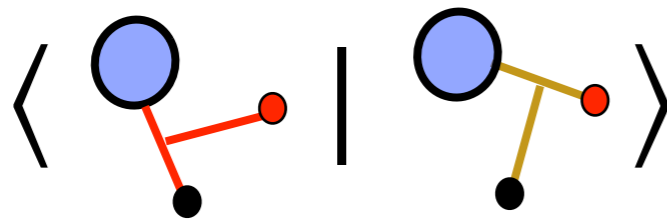
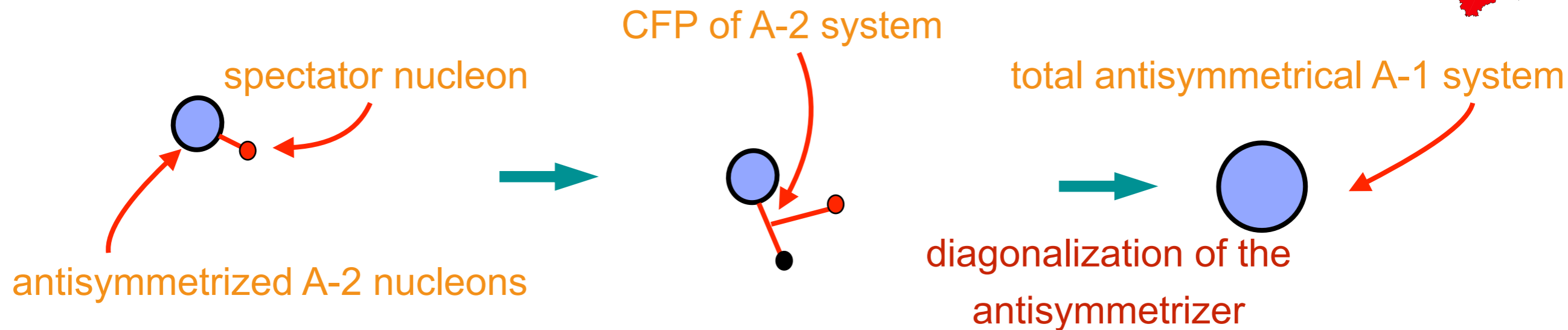
$$\langle \text{blue circle with red dot} | V_{YN} | \text{blue circle with red dot} \rangle = \langle \text{blue circle with red dot} | \text{blue circle with red dot and black dot} \rangle \langle \text{blue circle with red dot and black dot} | V_{YN} | \text{blue circle with red dot and black dot} \rangle \langle \text{blue circle with red dot and black dot} | \text{blue circle with red dot} \rangle$$

Application of to NN, YN, 3N and YNN interactions require the representation of particle transitions.

(see Liebig et al. EPJ A 52,103 (2016), Le et al. EPJ A 56, 301 (2020)  
for combinatorial factors see Le et al. EPJ A 57, 217 (2021))

# Jacobi-NCSM — CFP

First, generate **antisymmetrized states** for the A-1 nucleon system



antisymmetrizer is equivalent to coordinate trafo  
expression in terms of Talmi-Moshinsky brackets

(Navrátil et al. PRC 61,044001(2000))

The CFP coefficients  $\langle \text{blue circle, red dot} | \text{large blue circle} \rangle$  are obtained by diagonalization of the antisymmetrizer.

**HO** states guarantee:

- complete separation of antisymmetrized and other states
- **independence** of HO length/frequency

These coefficients will be openly accessible as **HDF5** data files

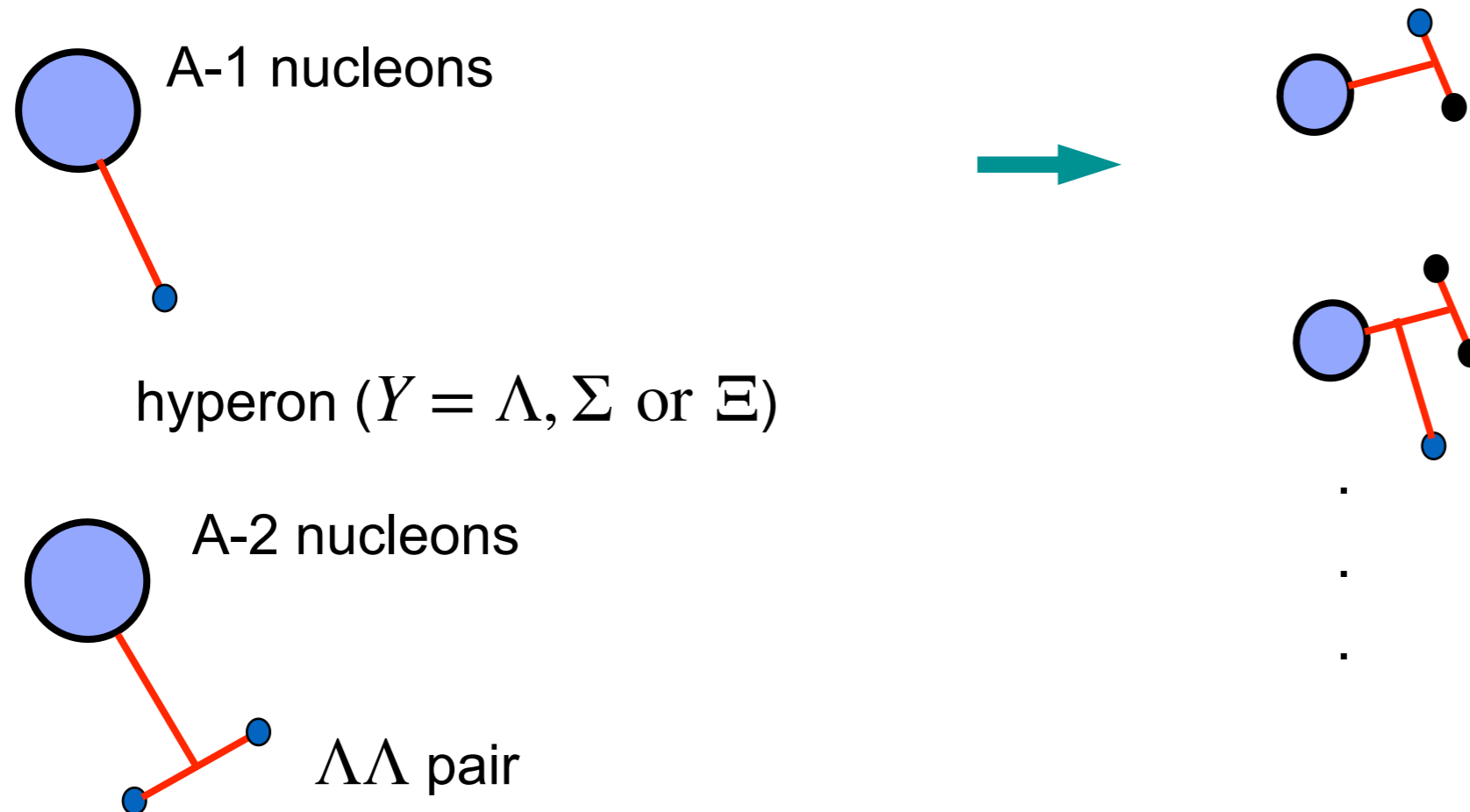
(download server is in preparation *(please contact me when interested!)*)

(Liebig et al. EPJ A 52,103 (2016))

# Jacobi-NCSM states for $S = -1$

**A-body hypernuclei state** (no antisymmetrization with respect to nucleons required)

Third, rearrange baryons for the application of interactions, ...



Again HO states guarantee the independence of HO length/frequency.

Transition coefficients are also accessible as **HDF5** data files to anyone interested.

(Le, Haidenbauer, Meißner, AN, 2020 & 2021)

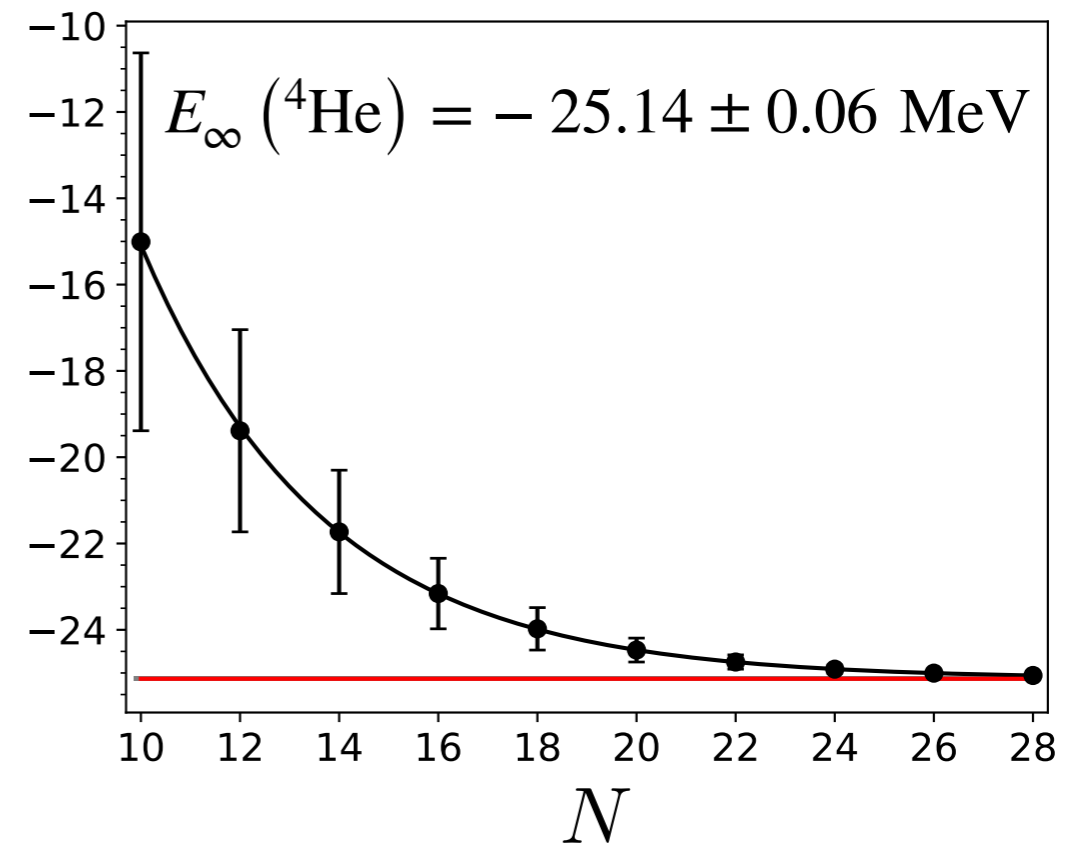
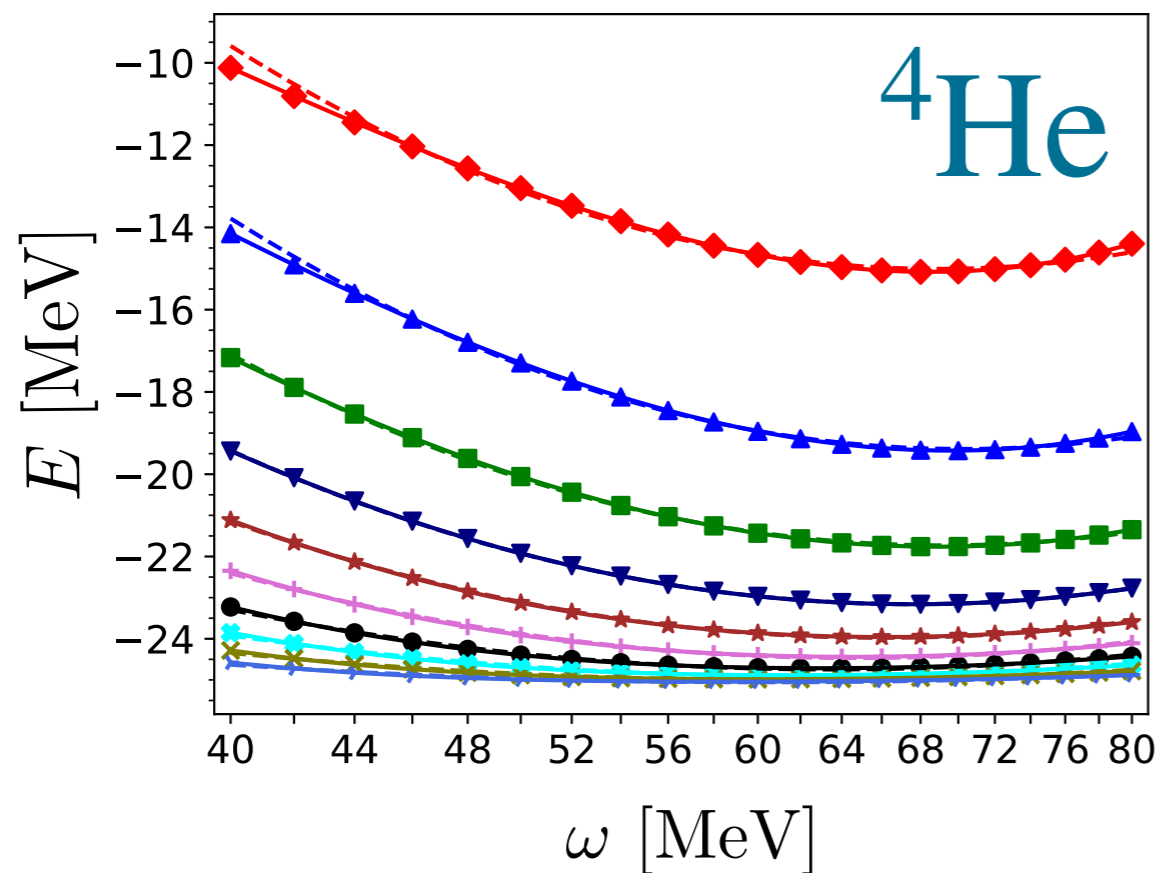
Converged results feasible for "**soft**" interactions.

# Convergence for Jacobi-NCSM

Simple example:  ${}^4\text{He}$  with SMS  $N^2\text{LO}(550)$

observed dependence on  $\omega$  and  $N$

$$E(\omega) = E_N + \kappa (\log(\omega) - \log(\omega_{opt}))^2 \longrightarrow E_N = E_\infty + A e^{-bN}$$



**Conservative** uncertainty estimate: difference of  $E_{N_{\max}}$  and  $E_\infty$   
Numerical uncertainties for light nuclei are small.

For p-shell, numerical uncertainty is more sizable due to smaller  $N_{\max}$   
and smaller separation energies. (Liebig et al. EPJ A 52,103 (2016))

In future: neural networks for extrapolation (see Wolfgruber et al. PRC 110,014327 (2024))



**Similarity renormalization group** is by now a **standard tool** to obtain soft effective interactions for various many-body approaches (NCSM, coupled-cluster, MBPT, ...)

Idea: perform a unitary transformation of the NN (and YN interaction) using a cleverly defined "generator"

(Bogner et al. PRC 75,061001 (2007))

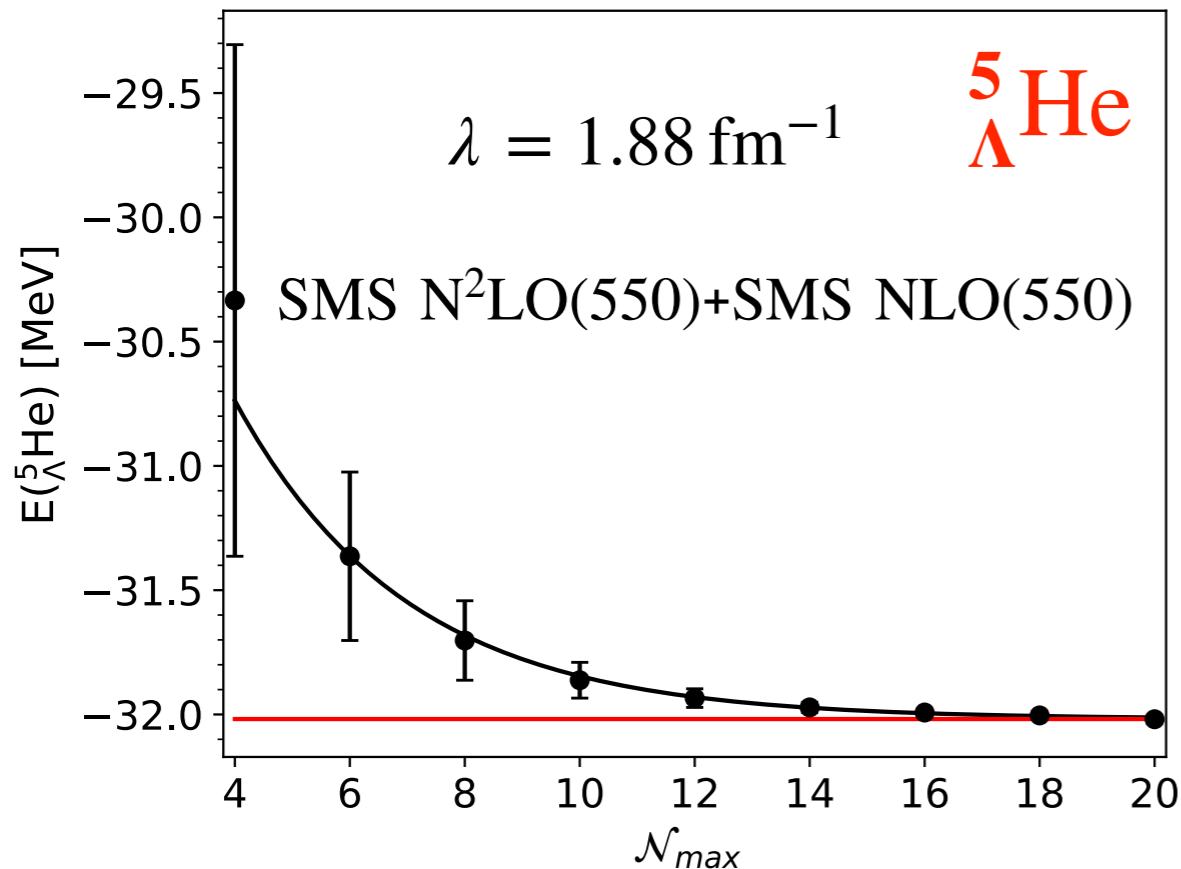
$$\frac{dH_s}{ds} = \left[ \underbrace{[T, H(s)]}_{\equiv \eta(s)}, H(s) \right] \quad H(s) = T + V(s)$$

this choice of generator drives  $V(s)$  into a diagonal form in momentum space

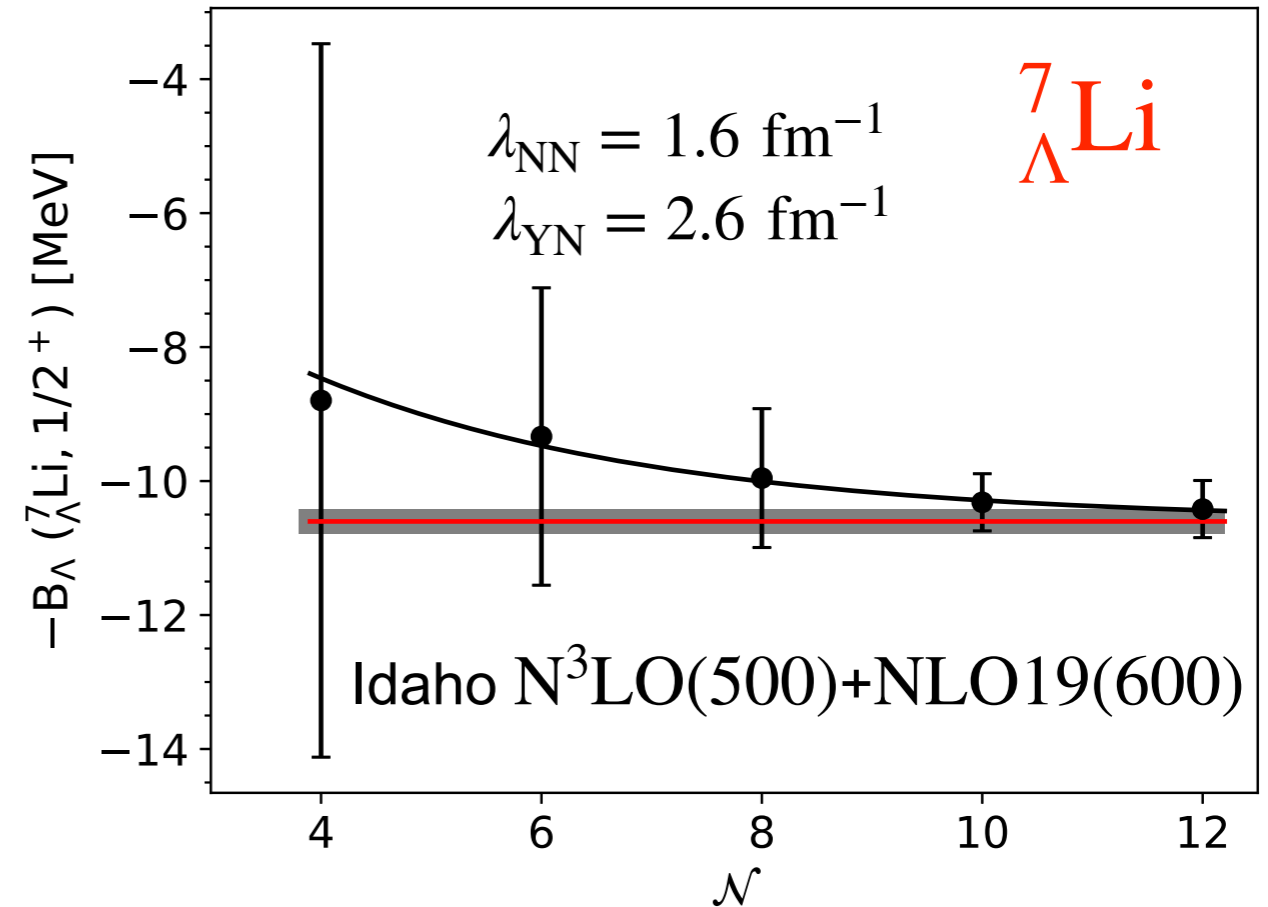
- $V(s)$  will be **phase equivalent** to original interaction
- short range  $V(s)$  will change towards **softer interactions**
- Evolution can be restricted to **2-,3-, ... body level** (approximation)
- $\lambda = \left( \frac{4\mu_{BN}^2}{s} \right)^{1/4}$  is a measure of the width of the interaction in momentum space
- **dependence** of results on  $\lambda$  or  $s$  is a measure for **missing terms**

# J-NCSM convergence

SRG evolution improves convergence



$$E({}^5_{\Lambda}\text{He}) = -32.018 \pm 0.001 \text{ MeV}$$



$$E_{\Lambda}({}^7_{\Lambda}\text{Li}) = 10.6 \pm 0.2 \text{ MeV}$$

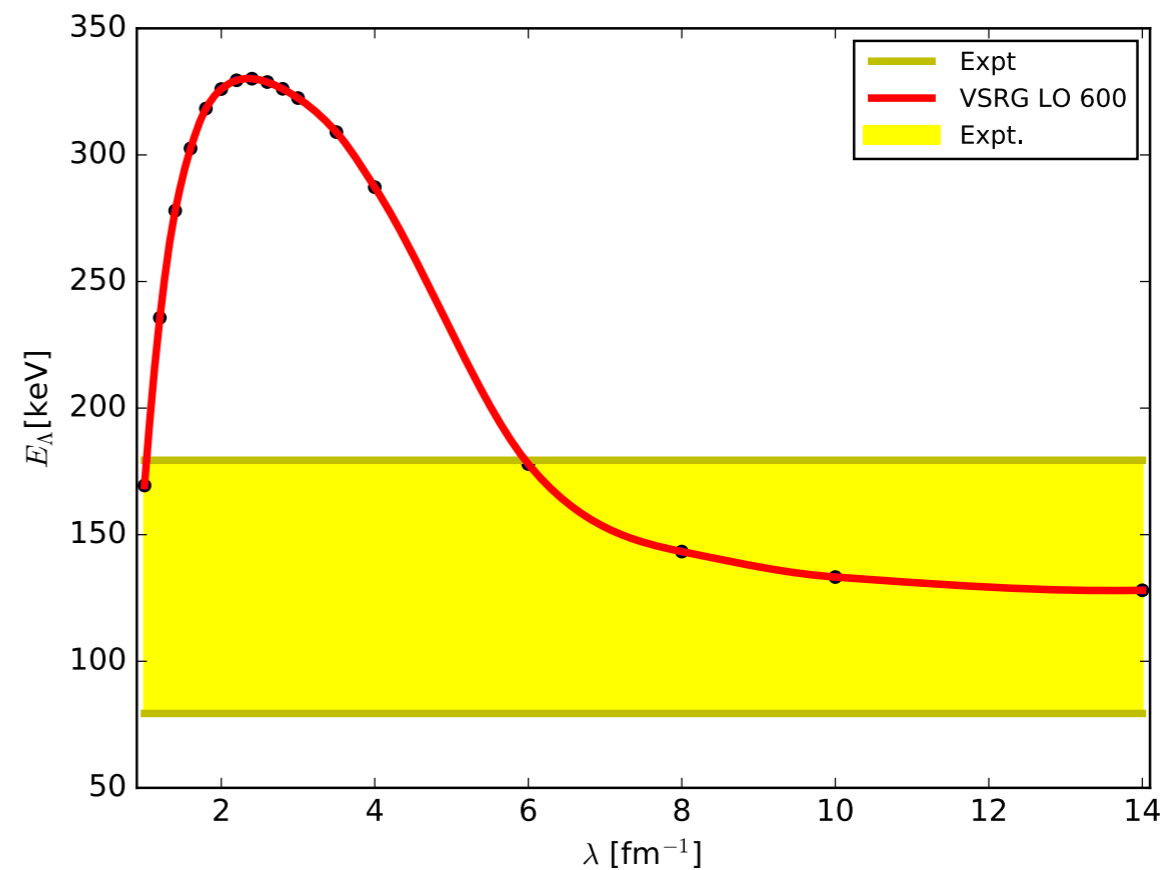
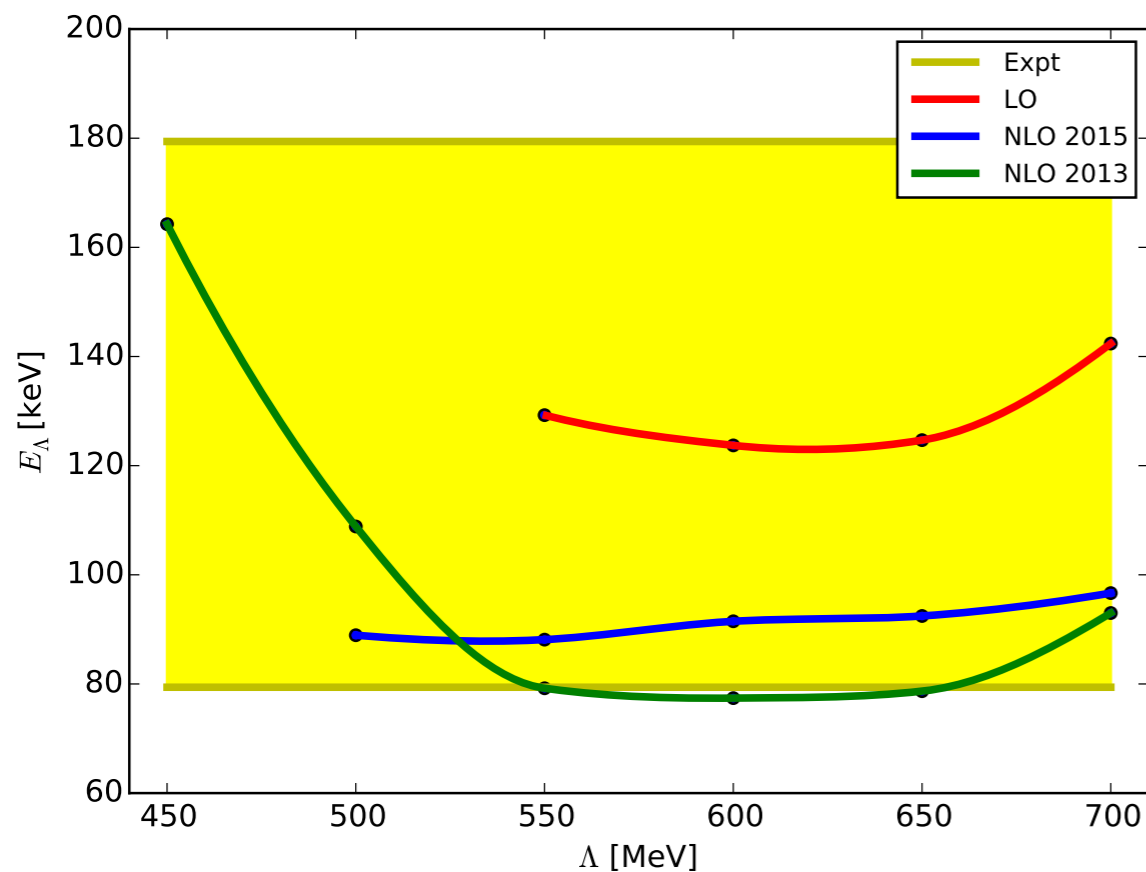
- for light nuclei and hypernuclei, the numerical uncertainty is negligible.
- for p-shell nuclei/hypernuclei, the uncertainty is visible
- extrapolation of separation energy can reduce uncertainty of this quantity

# Induced 3BF ...

SRG parameter dependence is significant when NN and YN interactions are evolved

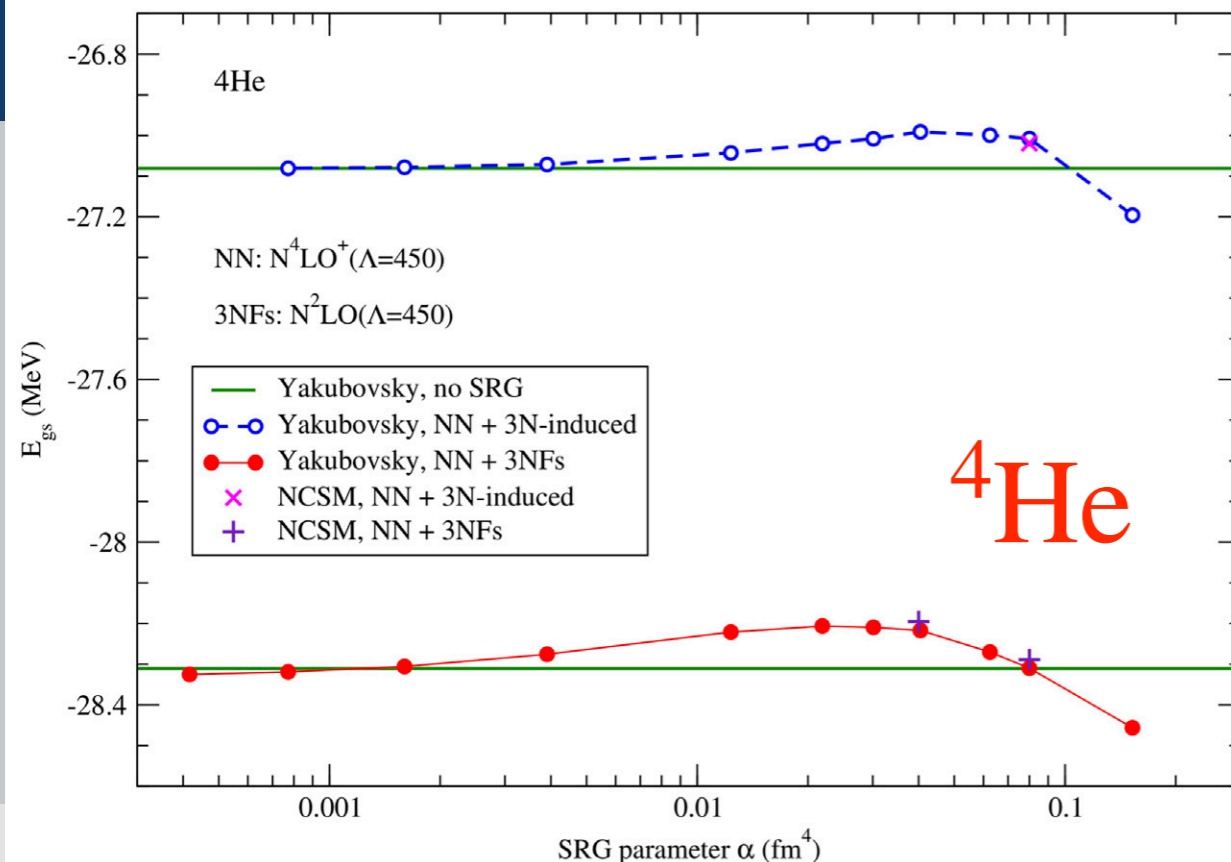
➡ missing 3N and YNN interactions

- 3NF is comparable to chiral 3NF
- YNN is larger than chiral YNN



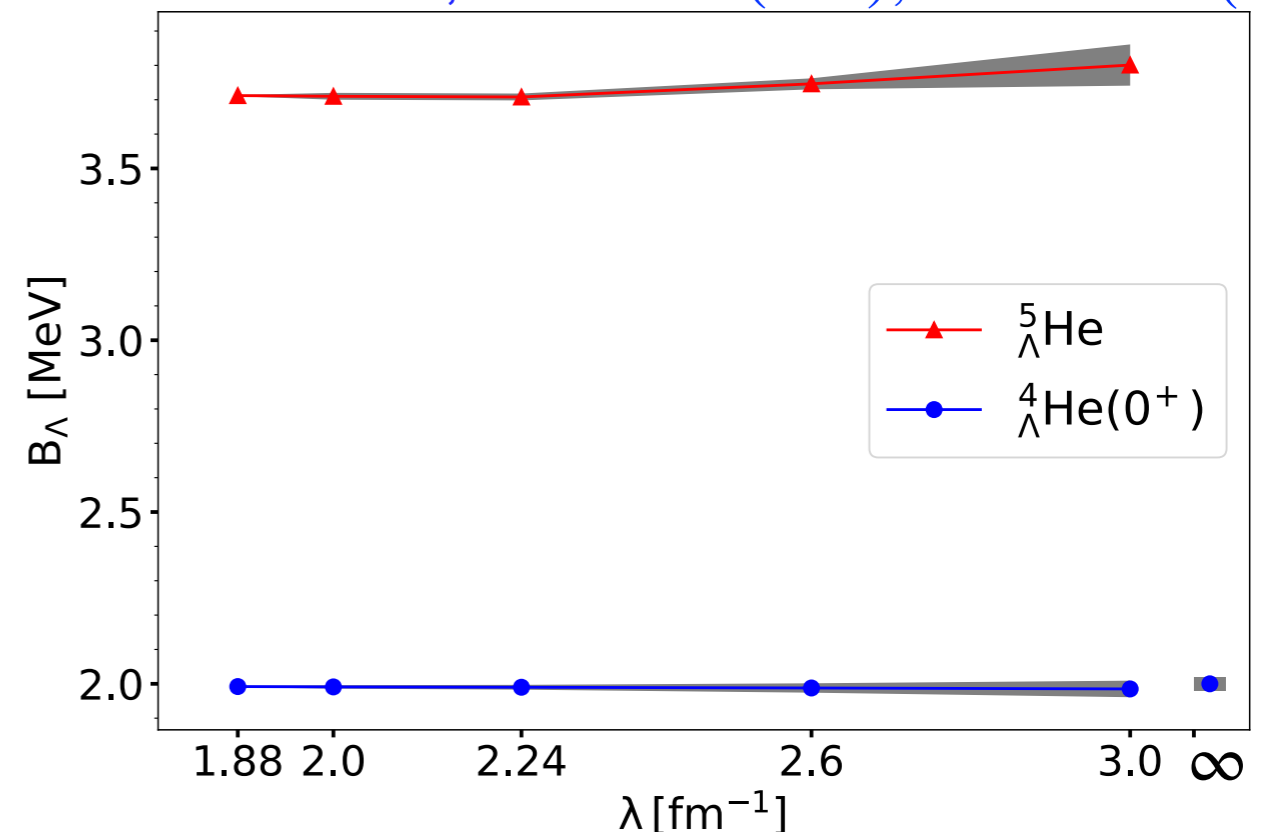
# SRG dependence of results

- SRG-induced 3N and YNN interactions
- $^4\text{He}$  binding energies varies by  $\approx 100 - 200$  keV (relevant in the future?)
- separation energies are even less dependent (YNNN forces small)



(Maris, Le, Nogga, Roth, Vary (2023))

NN:  $N^4\text{LO}^+$ , 3N:  $N^2\text{LO}(450)$ ; YN:  $N^2\text{LO}(550)$

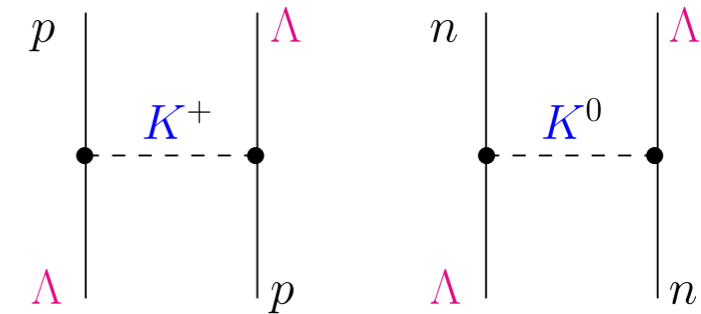


(Le (2023))

For **hypernuclei**, calculations based on SRG induced BB and 3B interactions are sufficiently accurate!

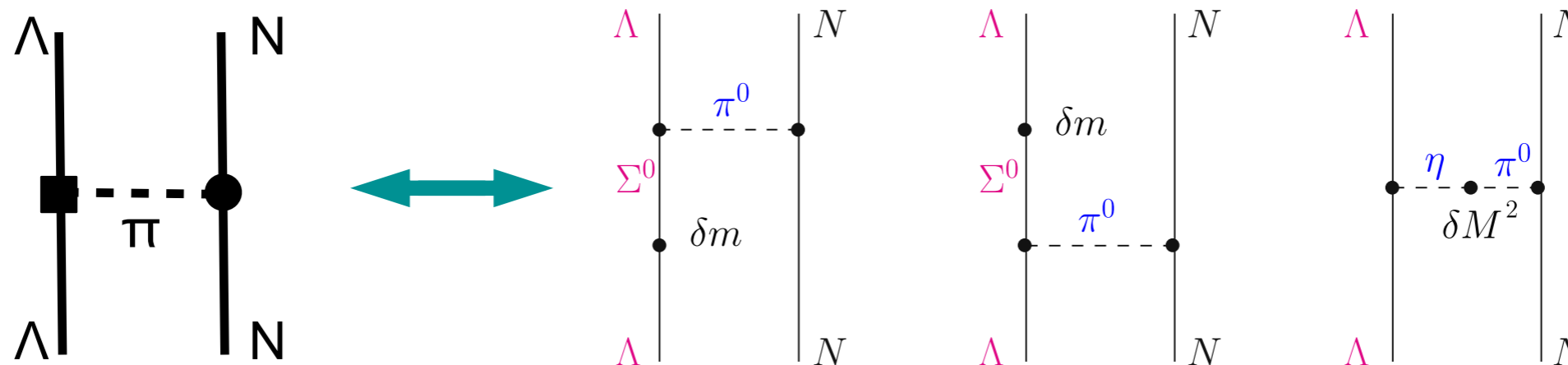
# CSB contributions to $\Lambda N$ interactions

- **formally leading** contributions:  
Goldstone boson mass difference
  - very small due to the small relative difference of kaon masses

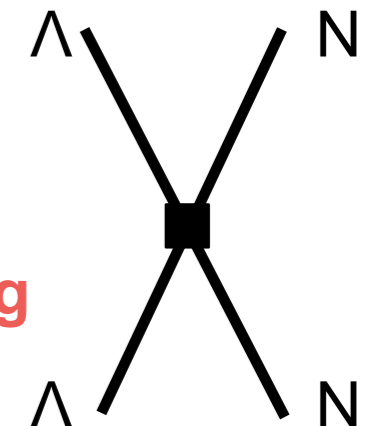


- **subleading but most important**
  - effective CSB  $\Lambda\Lambda\pi$  coupling constant (Dalitz, von Hippel, 1964)

$$f_{\Lambda\Lambda\pi} = \left[ -2 \frac{\langle \Sigma^0 | \delta m | \Lambda \rangle}{m_{\Sigma^0} - m_{\Lambda}} + \frac{\langle \pi^0 | \delta M^2 | \eta \rangle}{M_{\eta}^2 - M_{\pi^0}^2} \right] f_{\Lambda\Sigma\pi} \approx (-0.0297 - 0.0106) f_{\Lambda\Sigma\pi}$$

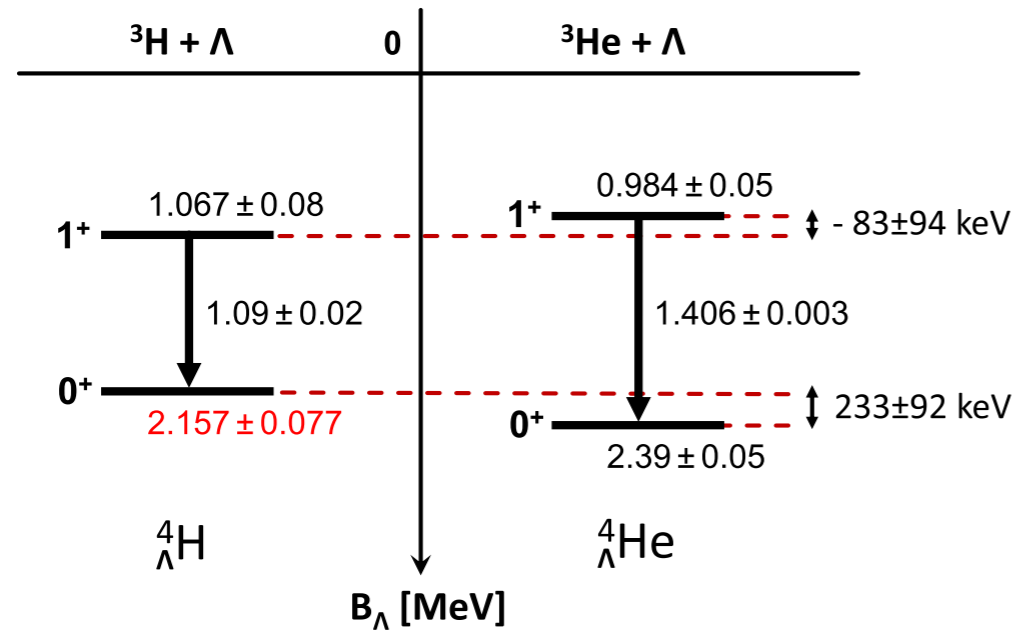


- **so far less considered but necessary for proper renormalization**
  - CSB contact interactions (for singlet and triplet)



**Aim: determine the two unknown CSB LECs and predict  $\Lambda n$  scattering**

# Fit of contact interactions



(Schulz et al., 2016; Yamamoto, 2015)

- Adjust the two CSB contact interactions to one main scenario (**CSB1**)
- update: Mainz average of CSB including new star data:  $178 \pm 55$  keV /  $-139 \pm 58$  keV

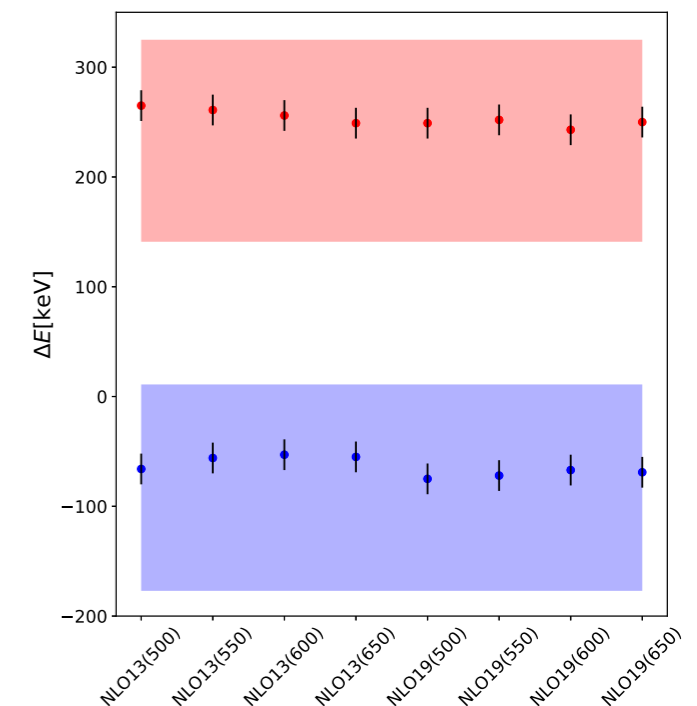
This was not used here.

- Fit of counter terms to data: size of LECs as expected by power counting

$$\frac{m_d - m_u}{m_u + m_d} \left( \frac{M_\pi}{\Lambda} \right)^2 C_{S,T} \approx 0.3 \cdot 0.04 \cdot 0.5 \cdot 10^4 \text{ GeV}^{-2} \propto 6 \cdot 10^{-3} \cdot 10^4 \text{ GeV}^{-2}$$

- Problem: large experimental uncertainty of experiment  
later adjust of CSB predictions is likely
- here only **fit to central values** to test theoretical uncertainties

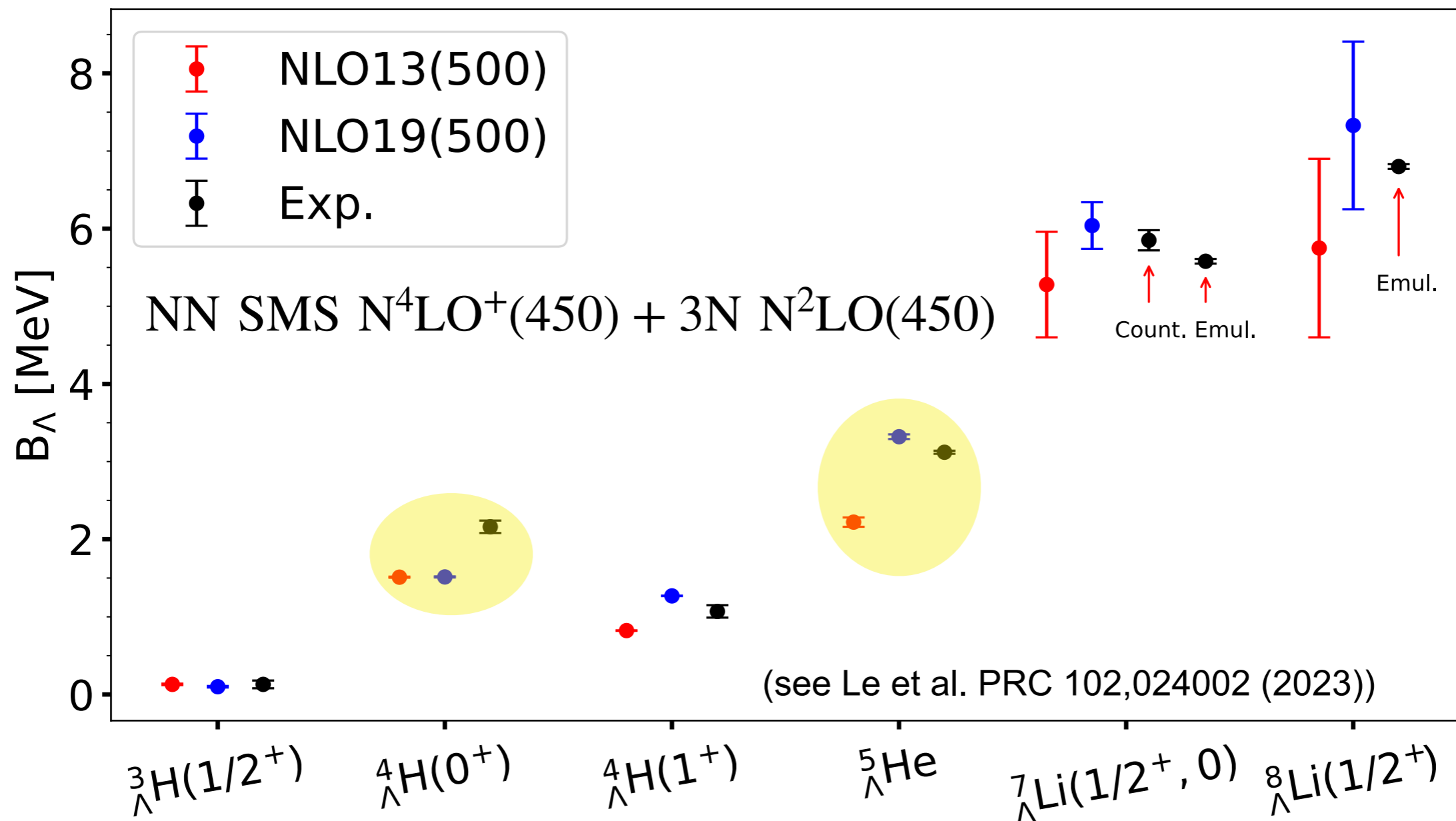
(see Haidenbauer et al. FBS 62,105 (2021))



# Application to $A = 7$ and 8

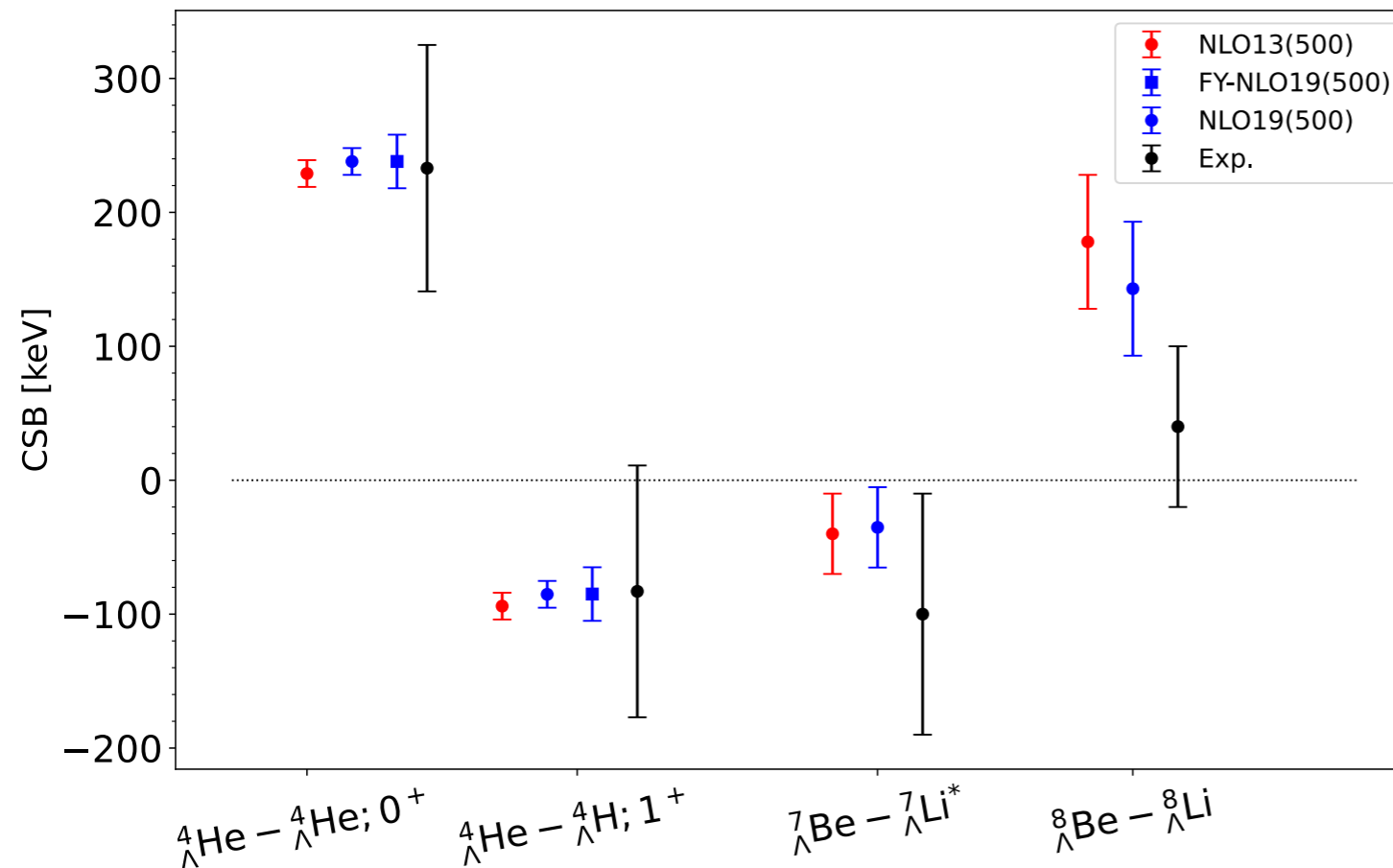


- YN interaction adjusted to the hypertriton — YNN is small
- based only on YN interactions: splitting for  ${}^4_{\Lambda}\text{H}$  is not well reproduced — YNN(?)
- NLO19 gives better results for  ${}^5_{\Lambda}\text{He}$  and heavier hypernuclei  
— accidentally small YNN interaction?
- uncertainties are numerical — no estimate of chiral uncertainties yet



# Predictions for $A = 2,7$ and 8

- CSB scattering length predicted **independent of the realization**
- keep in mind: CSB still not fixed — experimental uncertainty is large
- scenario studied here is only **marginally consistent** with CSB in  $A = 8$



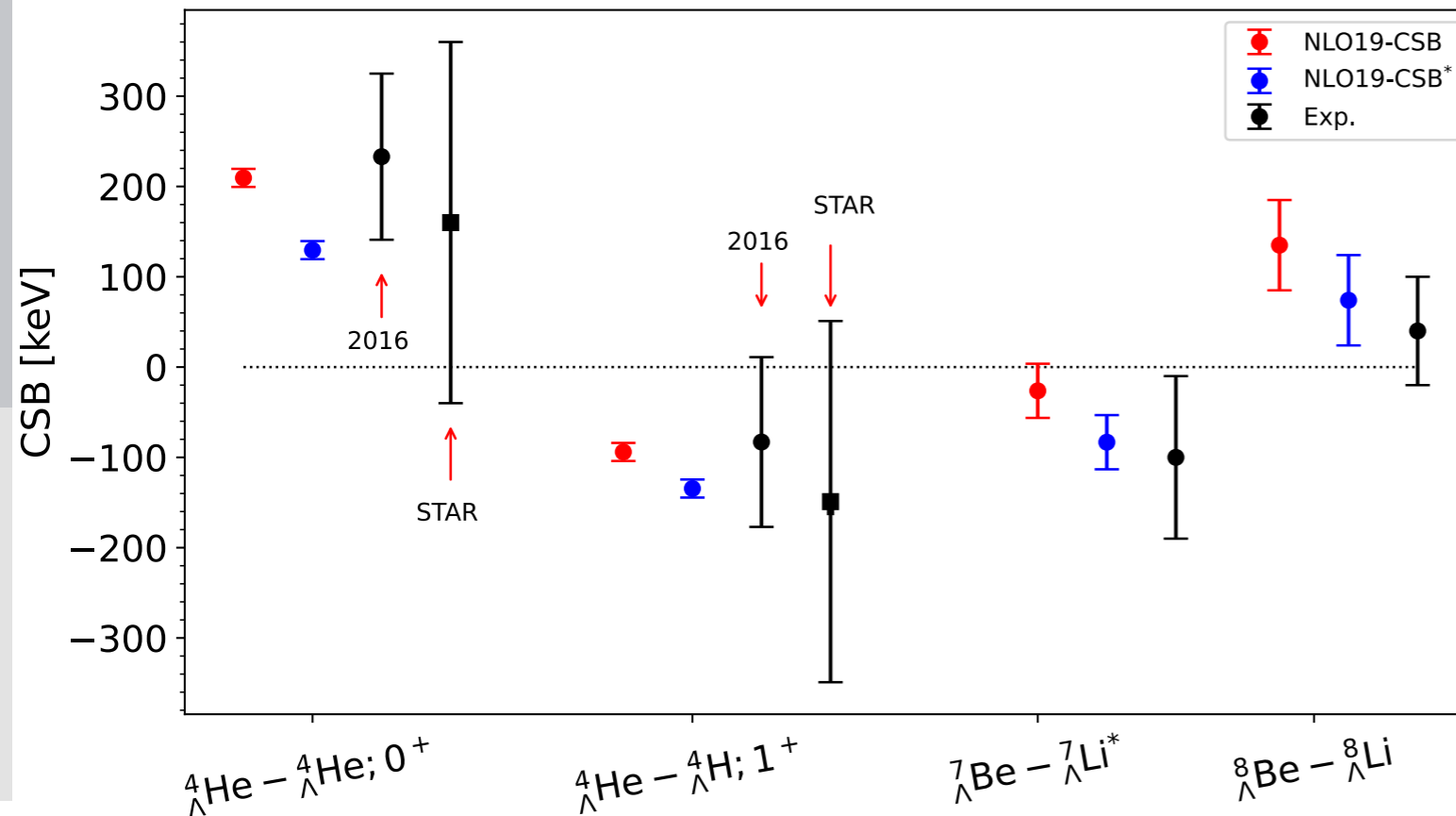
(Le et al. PRC 102,024002 (2023))

	$a_s^{\Lambda p}$	$a_t^{\Lambda p}$	$a_s^{\Lambda n}$	$a_t^{\Lambda n}$
NLO13(500)	-2.604	-1.647	-3.267	-1.561
NLO13(550)	-2.586	-1.551	-3.291	-1.469
NLO13(600)	-2.588	-1.573	-3.291	-1.487
NLO13(650)	-2.592	-1.538	-3.271	-1.452
NLO19(500)	-2.649	-1.580	-3.202	-1.467
NLO19(550)	-2.640	-1.524	-3.205	-1.407
NLO19(600)	-2.632	-1.473	-3.227	-1.362
NLO19(650)	-2.620	-1.464	-3.225	-1.365

(Haidenbauer et al. FBS 62,105 (2021))

# New STAR data for $A = 4$ CSB

- fit to STAR data only
- only slight adjustment required
- improves description to p-shell CSB
- higher experimental accuracy is desirable
- good example of using hypernuclei to determine YN interactions



	NLO19(500)	CSB	CSB*
$a_s^{\Lambda p}$	-2.91	-2.65	-2.58
$a_s^{\Lambda n}$	-2.91	-3.20	-3.29
$\delta a_s$	0	0.55	0.71
$a_t^{\Lambda p}$	-1.42	-1.57	-1.52
$a_t^{\Lambda n}$	-1.41	-1.45	-1.49
$\delta a_t$	-0.01	-0.12	-0.03

(see Le et al. PRC 102,024002 (2023))

# Uncertainty analysis to $A = 3$ to 5

Order N<sup>2</sup>LO requires combination of chiral NN, YN, 3N and **YNN** interaction

Results for **different orders** enable uncertainty estimate:

Ansatz for the order by order convergence:

$$X_K = X_{ref} \sum_{k=0}^K c_k Q^k \quad \text{where} \quad Q = M_{\pi}^{eff} / \Lambda_b \quad (X_{ref} \text{ LO, exp., max, ...})$$

**Bayesian analysis** of the uncertainty following Melendez et al. 2017,2019

**Extracting  $c_k$  for  $k \leq K$  from calculations**

➡ **probability distributions for  $c_k$**

➡ 
$$\delta X_K = X_{ref} \sum_{k=K+1}^{\infty} c_k Q^k$$

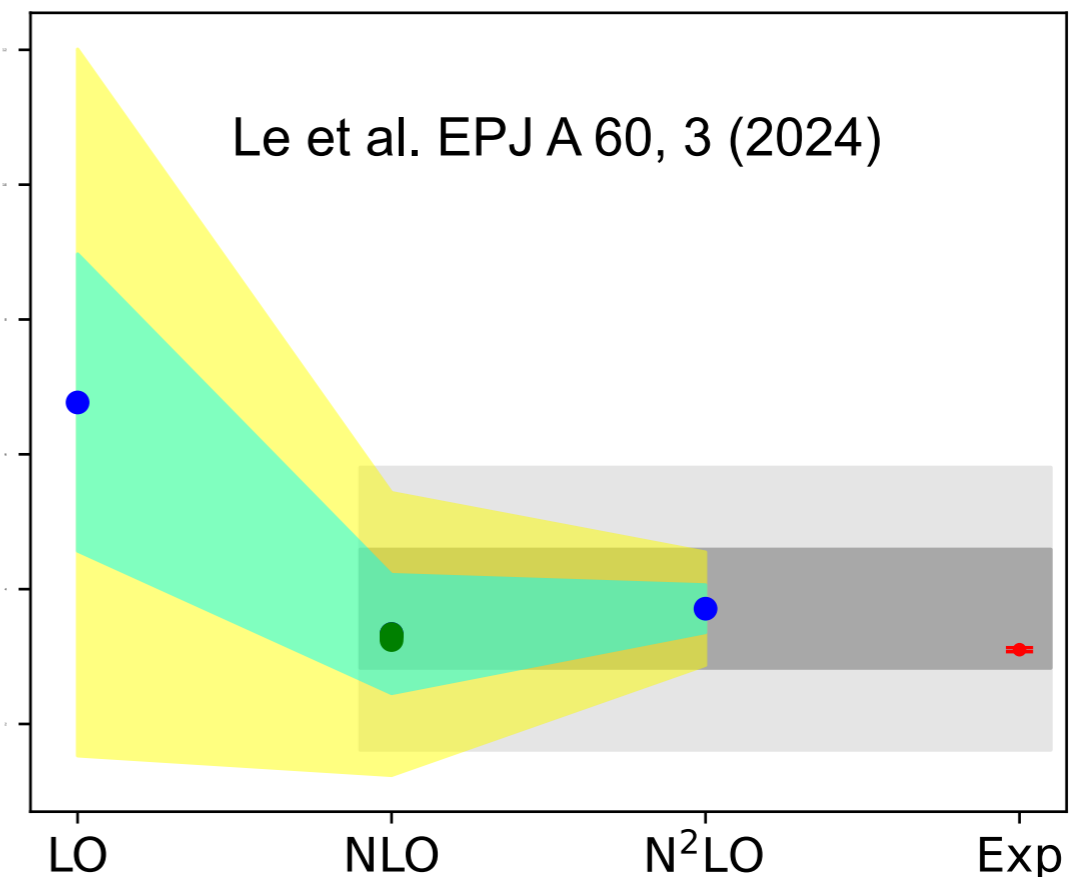
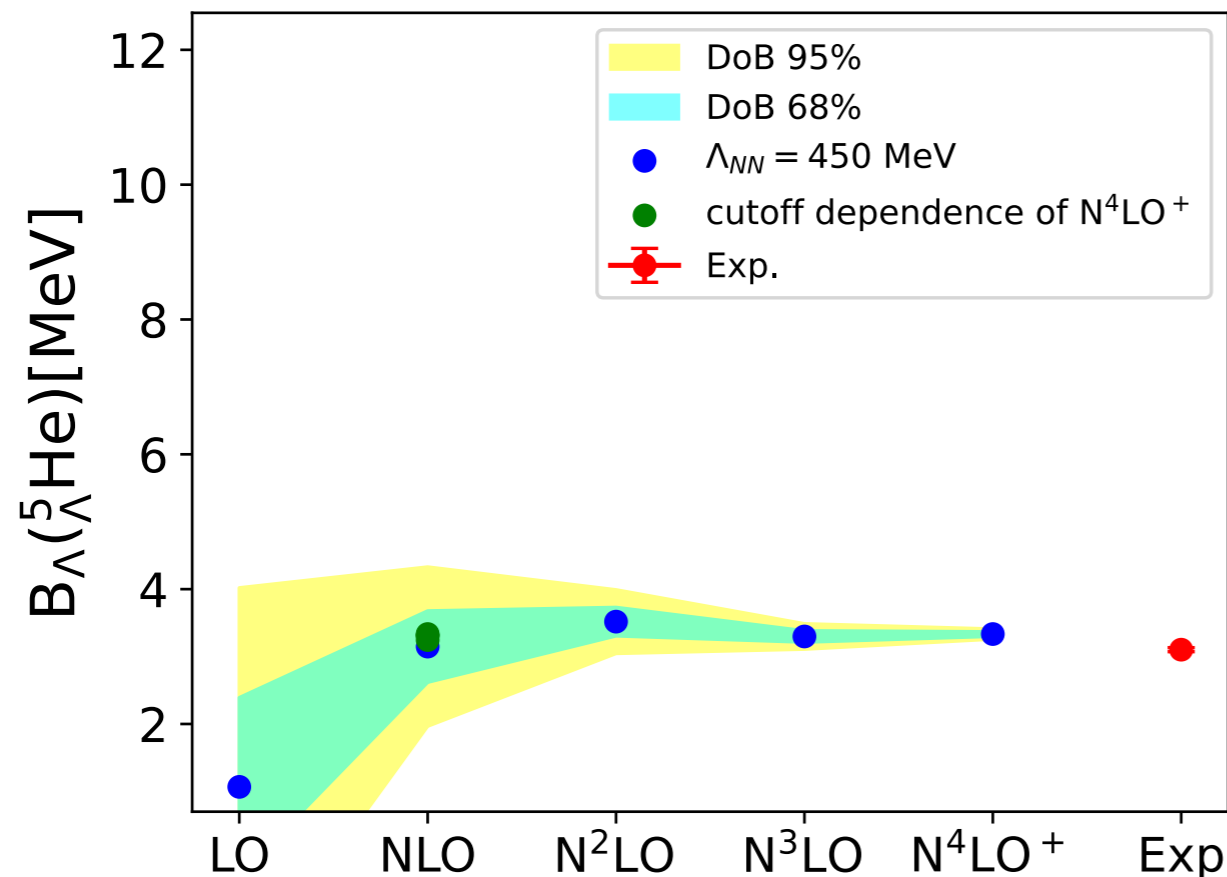
**Uncertainty due to missing higher orders is more relevant  
than numerical uncertainty! (for light nuclei)**

# Application to ${}^5_{\Lambda}\text{He}$ and summary

- **without YNN**: sizable uncertainties at  $A = 4$  and 5
- $A = 3$  sufficiently accurate
- NN/YN dependence small at least for  $A = 3$

nucleus	$\Delta_{68}(NN)$	$\Delta_{68}(YN)$
${}^3_{\Lambda}\text{H}$	0.011	0.015
${}^4_{\Lambda}\text{He} (0^+)$	0.157	0.239
${}^4_{\Lambda}\text{He} (1^+)$	0.114	0.214
${}^5_{\Lambda}\text{He}$	0.529	0.881

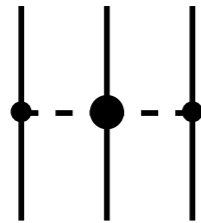
→ at the same time: estimate of YNN !



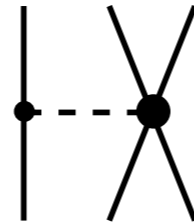
# YNN ( $\Lambda$ NN) interactions

Leading 3BF with the usual topologies (Petschauer et al. PRC 93, 014001 (2016))

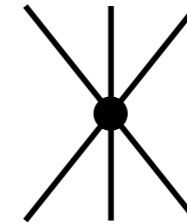
ChPT  $\longrightarrow$  all octet mesons contribute  $\longrightarrow$  **only take  $\pi$  explicitly into account**



2 LECs in  $\Lambda$ NN  
(up to 10)



2 LECs in  $\Lambda$ NN  
(up to 14)

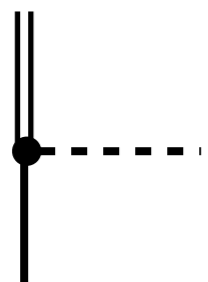


3 LECs in  $\Lambda$ NN  
5 LECs in  $\Sigma$ NN + 1  $\Lambda$ - $\Sigma$  transition

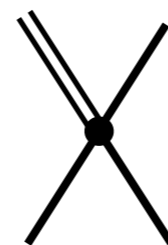
only few data  $\longrightarrow$  need to keep the **# of LECs** small  
Decuplet baryons ( $\Sigma^*$ ...) might enhance YNN partly to NLO

(Petschauer et al., NPA 957, 347 (2017))

By decuplet saturation all LECs can be related to the following  
leading octet-decuplet transitions (Petschauer et al. Front. Phys. 8,12 (2020))



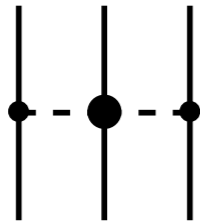
$$\propto C = \frac{3}{4}g_A$$



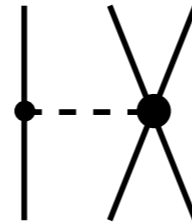
$\propto G_1, G_2 \longrightarrow$  **reduction to 2 LECs**

# YNN ( $\Lambda$ NN) interactions

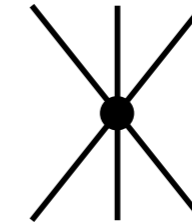
Decuplet saturation relates all LECs to  $G_1$  and  $G_2$



$$\propto C^2$$



$$\propto CG_1, CG_2$$



$$\propto (G_1)^2, (G_2)^2, G_1 G_2$$

For  $\Lambda$ NN:  $\propto C^2$

$$\propto C(G_1 + 3G_2)$$

$$\propto (G_1 + 3G_2)^2 \quad \text{1 LEC}$$

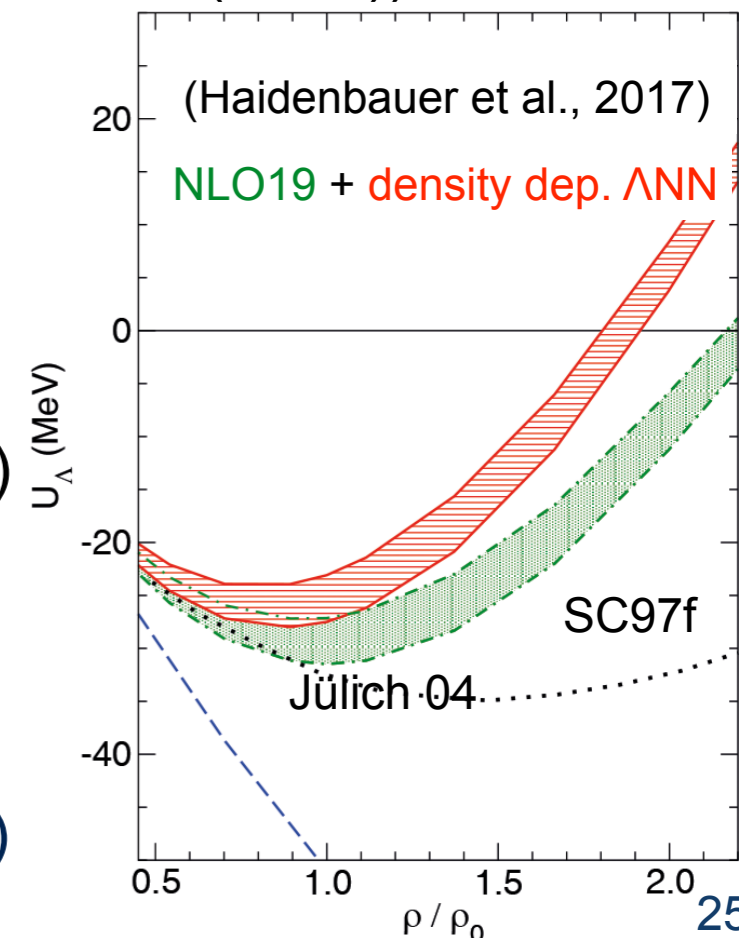
- ➡ density dependent BB interactions (Petschauer et al., NPA 957, 347 (2017))
- ➡ application to nuclear matter (Haidenbauer et al., EPJ A 53, 121 (2017))
- neutron stars (Logoteta et al., EJA 55, 207 (2019))

- contribution on the single particle potentials can be large
- realistic results seem to require partly cancelations of  $2\pi$  and  $1\pi$  exchange

(fixes sign of  $G_1 + 3G_2$ !)

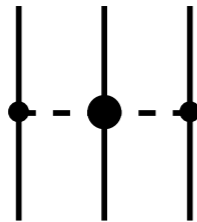
Recently: successful benchmark of matrix elements  
(Hoai Le et al. EPJ A 61,21 (2025))

and first direct application to light hypernuclei including  $\Sigma$ 's  
(Hoai Le et al. PRL 134, 072502 (2025))

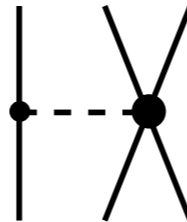


# YNN ( $\Lambda$ NN) interactions in practice

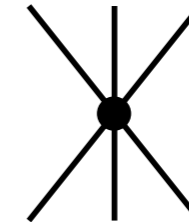
## Decuplet approximation in YNN



$$\propto C^2$$



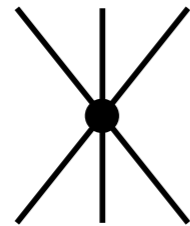
$$\propto CG_1, CG_2$$



$$\propto (G_1)^2, (G_2)^2, G_1 G_2$$

is not sufficient to fix spin dependence

➡ +  $\Lambda$ NN contact terms **without decuplet constraints**



$$\Lambda\text{NN} \propto C'_1, C'_2, C'_3$$

**ad hoc choice:** alter  $C_2$ :

$$C'_1 = C'_3 = \frac{(G_1 + 3G_2)^2}{72\Delta}$$

$$C'_2 = 0$$

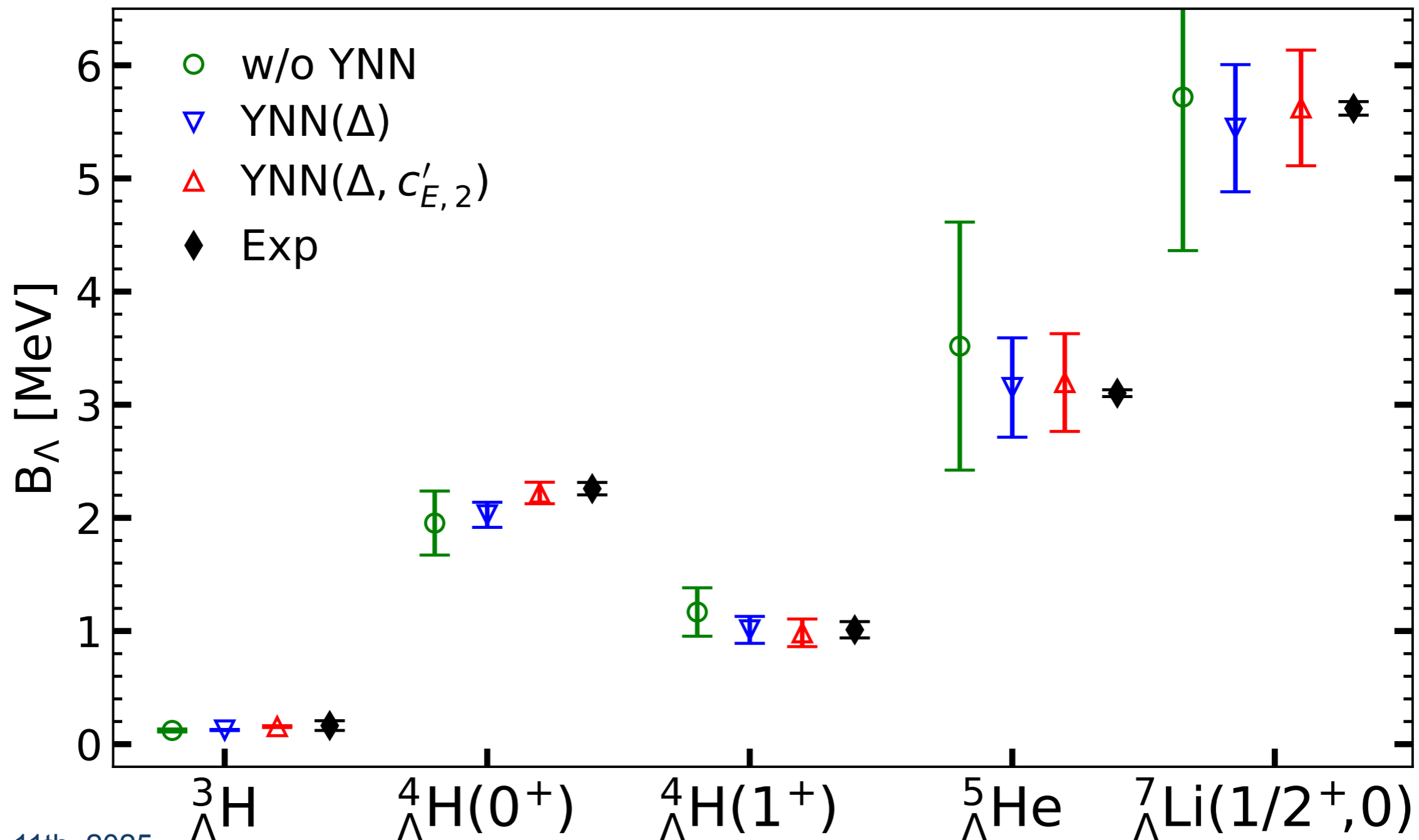


$$V_{\Lambda\text{NN}} = C'_2 \vec{\sigma}_1 \cdot (\vec{\sigma}_2 + \vec{\sigma}_3) (1 - \vec{\tau}_2 \cdot \vec{\tau}_3)$$

$$C'_2 = G_3$$

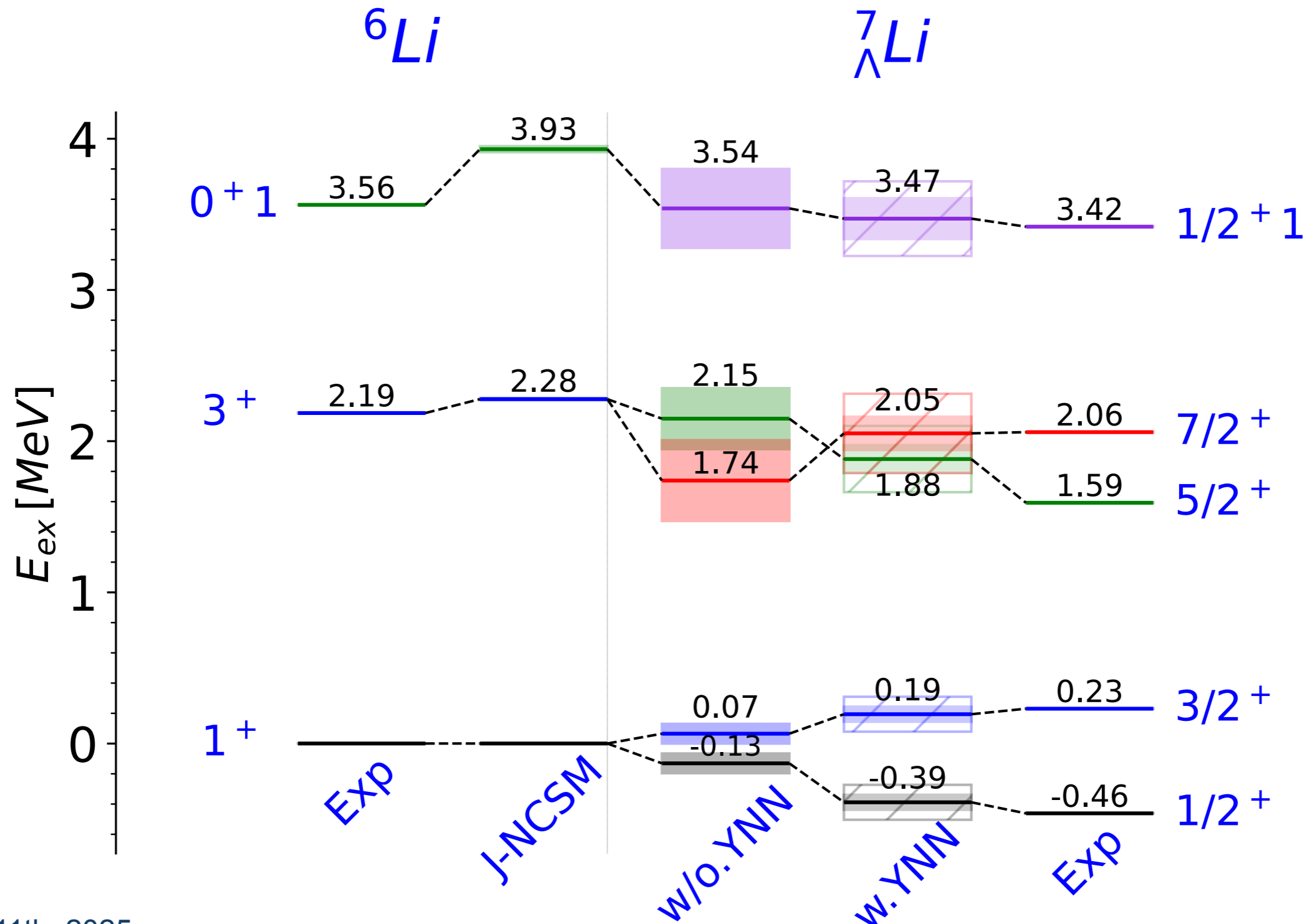
$C'_2$  introduces a spin dependent interaction in the most relevant particle channel

- Fit to  $0^+$  and  $1^+$  state of  ${}^4_{\Lambda}\text{He}$  and/or  ${}^5_{\Lambda}\text{He}$
- spin-dependence in  $A=4$  not well explained by decuplet saturation
- $C'_2$  term improves  $0^+$  of  ${}^4_{\Lambda}\text{He}$  and  $1/2^+$  of  ${}^7_{\Lambda}\text{Li}$
- agreement generally much better than  $N^2\text{LO}$  uncertainty



# YNN prediction for ${}^7_{\Lambda}\text{Li}$

- good agreement
- $C'_2$  term included, but not very important (not shown)
- higher states have significant uncertainty



# $S = -2$ hypernuclei — ${}_{\Lambda\Lambda}^6\text{He}$

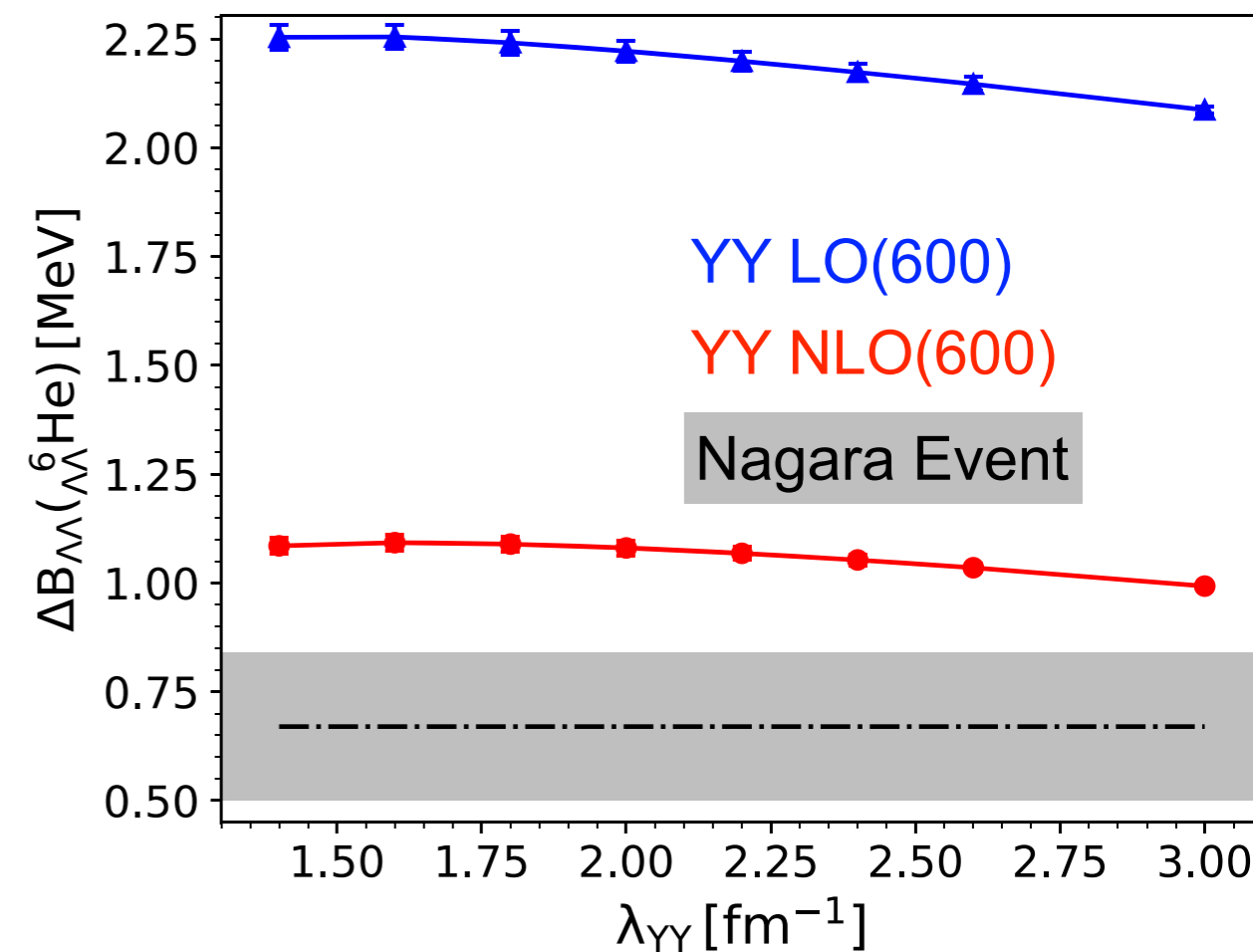
- $\Lambda\Lambda$  excess binding energy

$$\begin{aligned}\Delta B_{\Lambda\Lambda} &= B_{\Lambda\Lambda} - 2B_{\Lambda} \\ &= 2E({}^A_{\Lambda}X) - E({}_{\Lambda\Lambda}^AX) - E({}^{A-2}X)\end{aligned}$$

- NN, YN and YY interactions contribute
- use NN and YN that describe nuclei and single  $\Lambda$  hypernuclei
- small  $\lambda_{YY}$  dependence
- LO overbinds YY
- NLO predicts binding fairly well

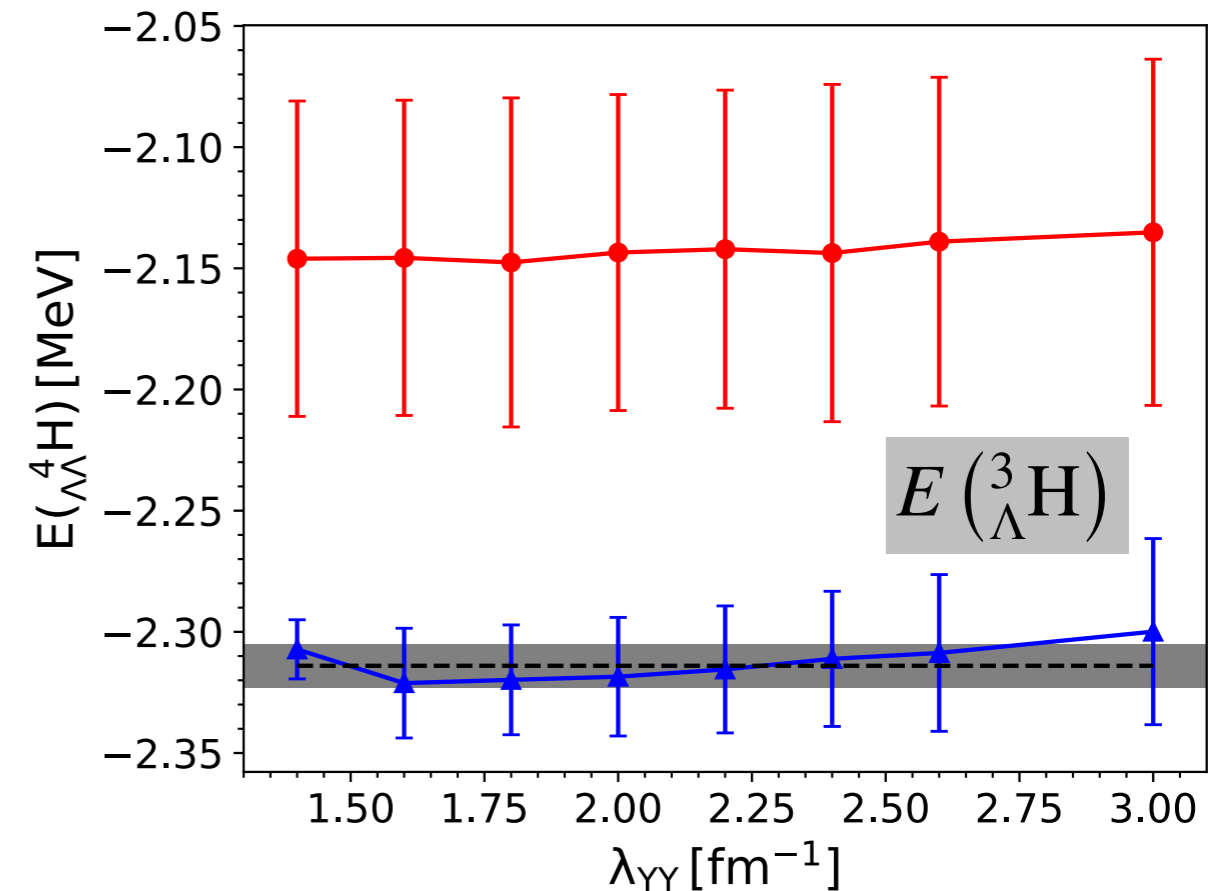
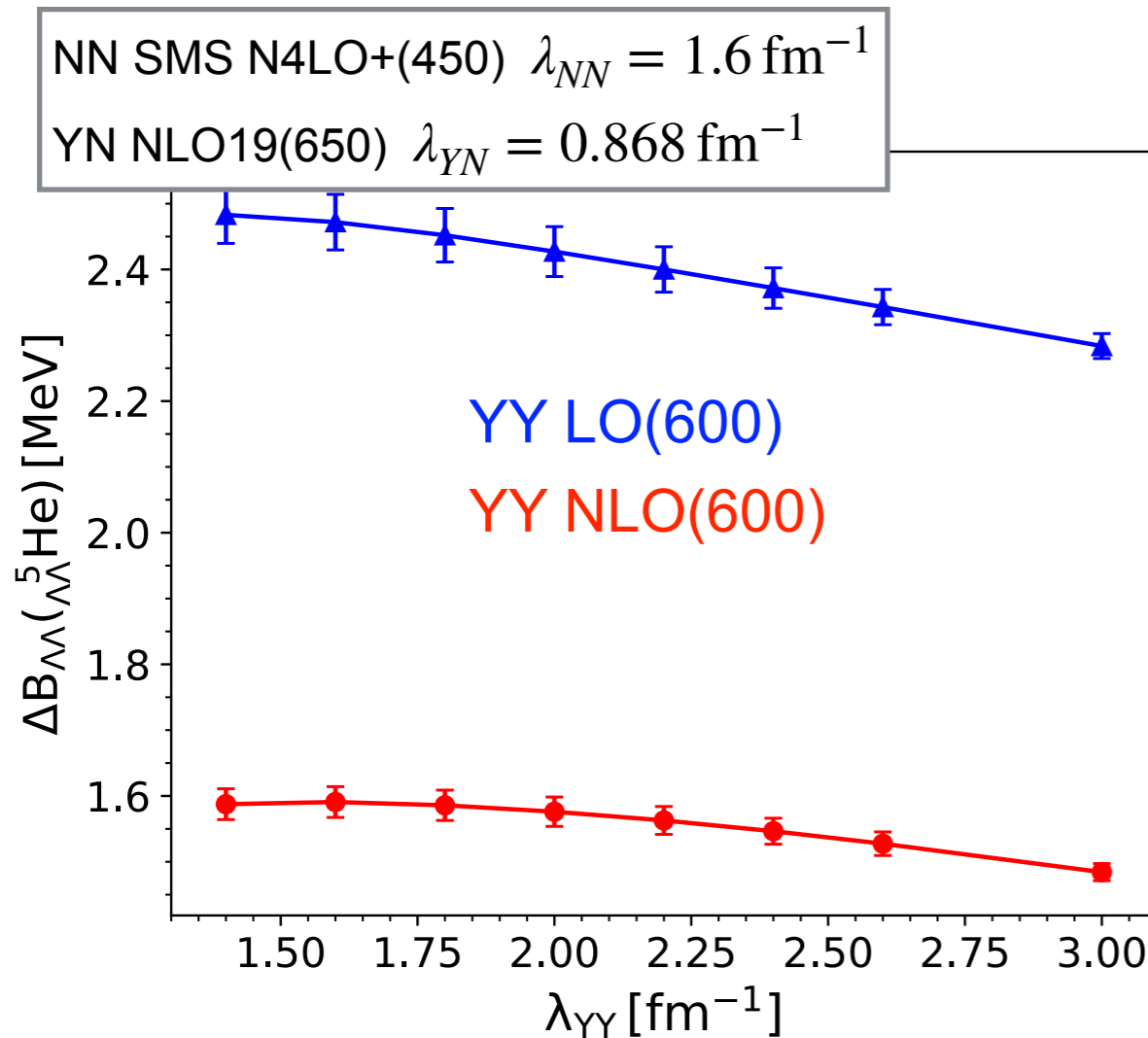
**Can an  $S = -2$  bound state for  $A = 4,5$  be expected?**

(Le et al., 2021)



NN SMS N4LO+(450)  $\lambda_{NN} = 1.6 \text{ fm}^{-1}$   
YN NLO19(650)  $\lambda_{YN} = 0.868 \text{ fm}^{-1}$

# $S = -2$ hypernuclei — ${}_{\Lambda\Lambda}^5\text{He}$ & ${}_{\Lambda\Lambda}^4\text{H}$



- $A = 5$ :  $\Lambda\Lambda$  excess binding energy &  $A = 4$ : binding energy
- $A = 5$ : LO & NLO predicts bound state
- $A = 4$ : NLO unbound, LO at threshold to binding (see also Contessi et al., 2019)
- excess energy larger for  $A = 5$  than for  $A = 6$  (in contrast to Filikhin et al., 2002!)

$S = -2$  bound state for  $A = 5$  can be expected,

for  $A = 4$  less likely but not ruled out!

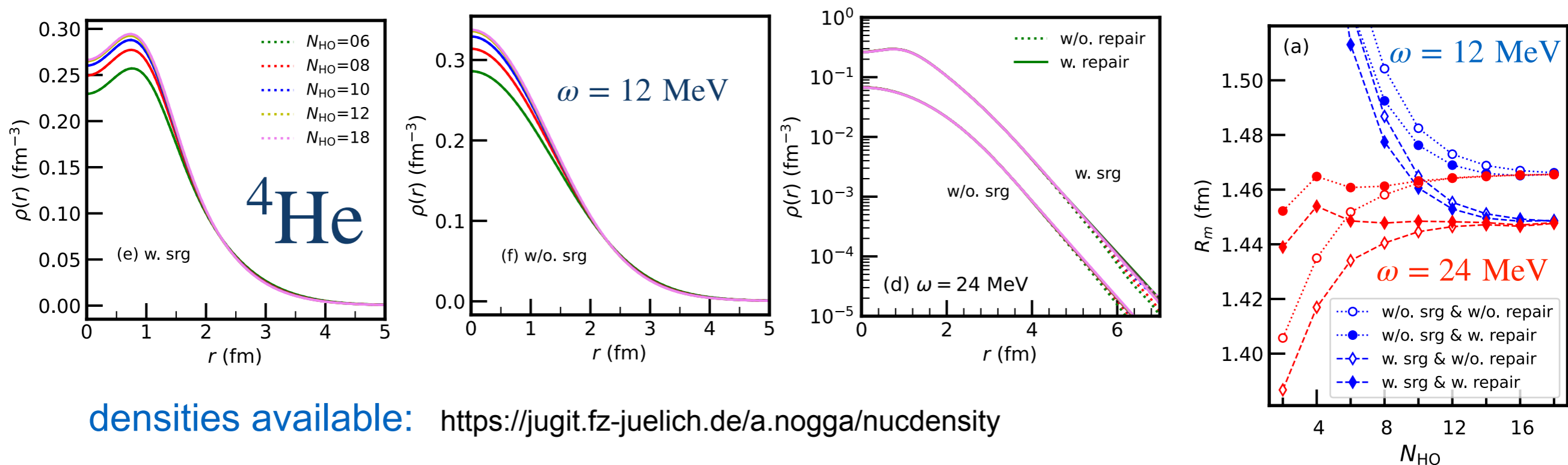
# Two-nucleon densities

- SRG evolution affects wave functions
- short-range, medium and high momentum observables affected
- unitary transformation of operators
- HO basis inefficient for describing exponential tail of wave functions
- HO frequency usually optimized for describing wave functions in range of interactions



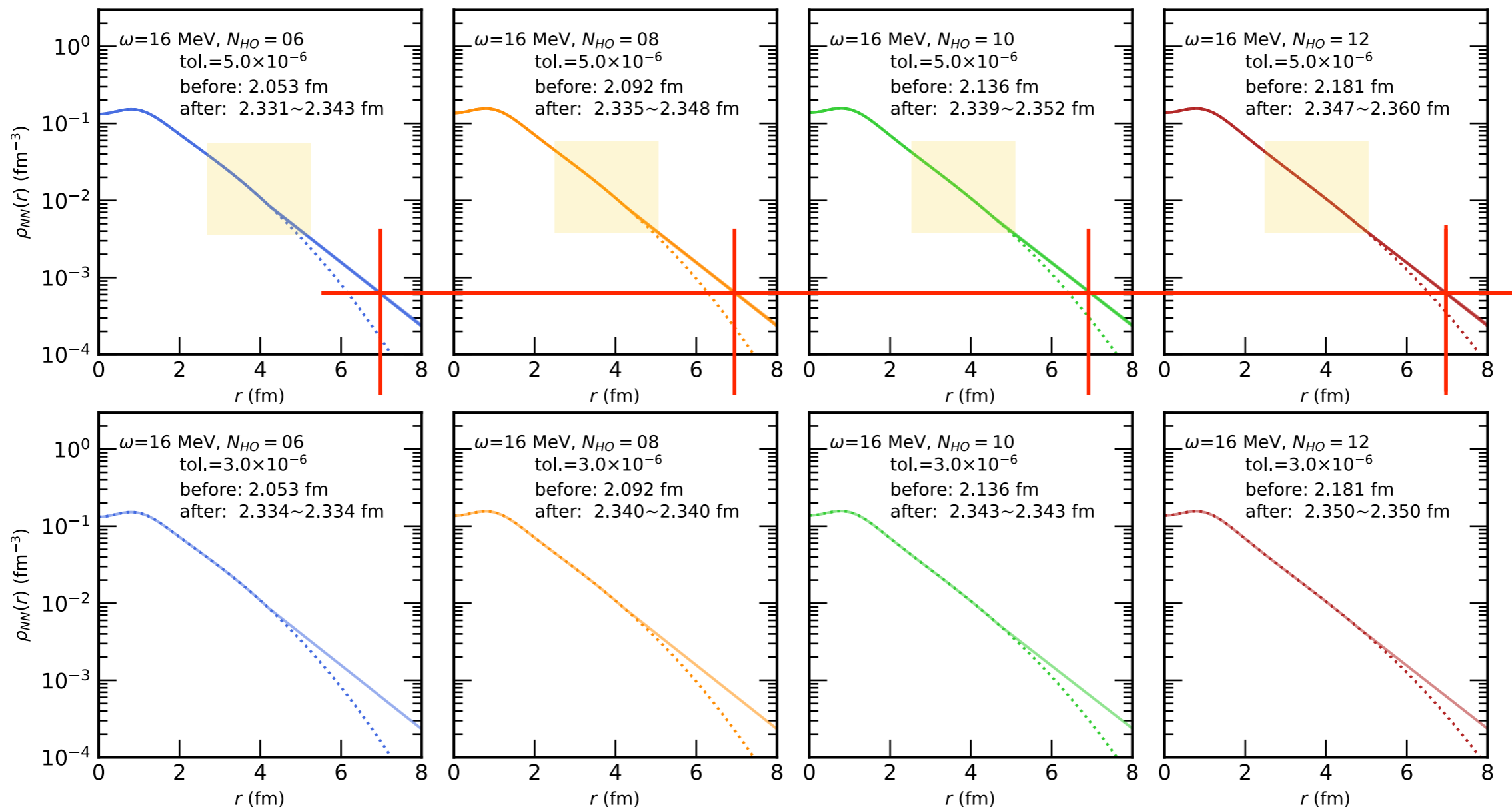
- define densities (in p- and r-space, 1-nucleon and 2-nucleon)
- apply SRG on densities (2-nucleon only!)
- correct long-range tail for long-ranged observables (2-nucleon only)
- calculations of matter and charge radii of light nuclei

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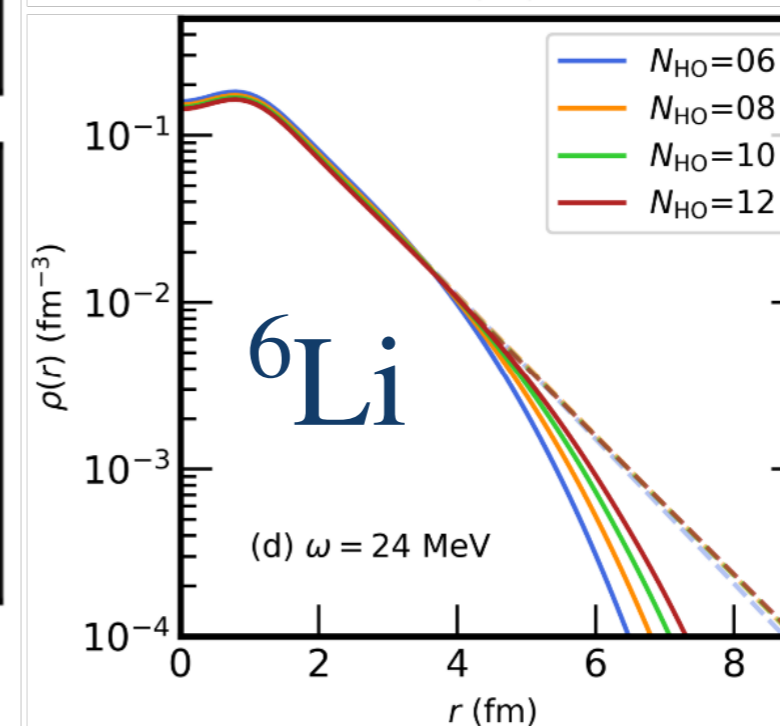
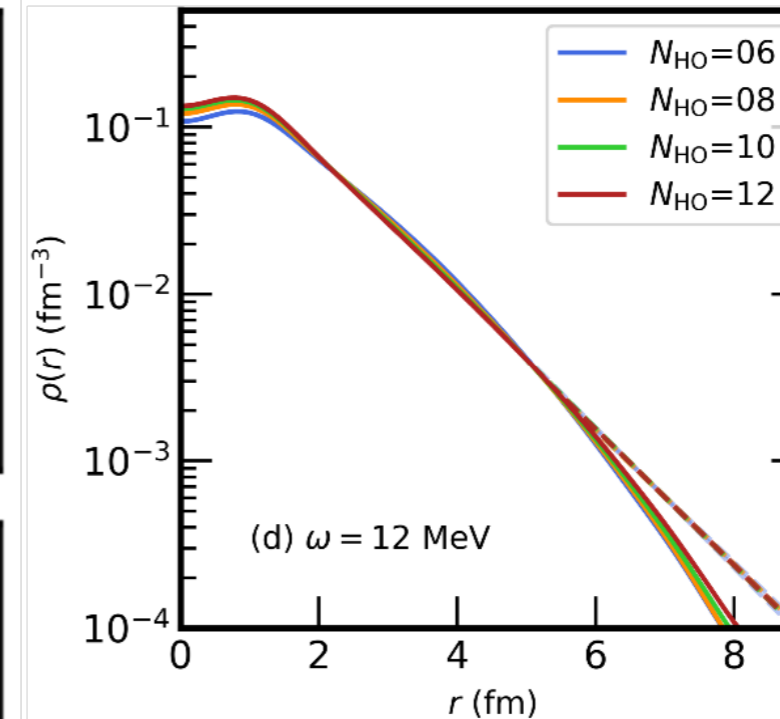
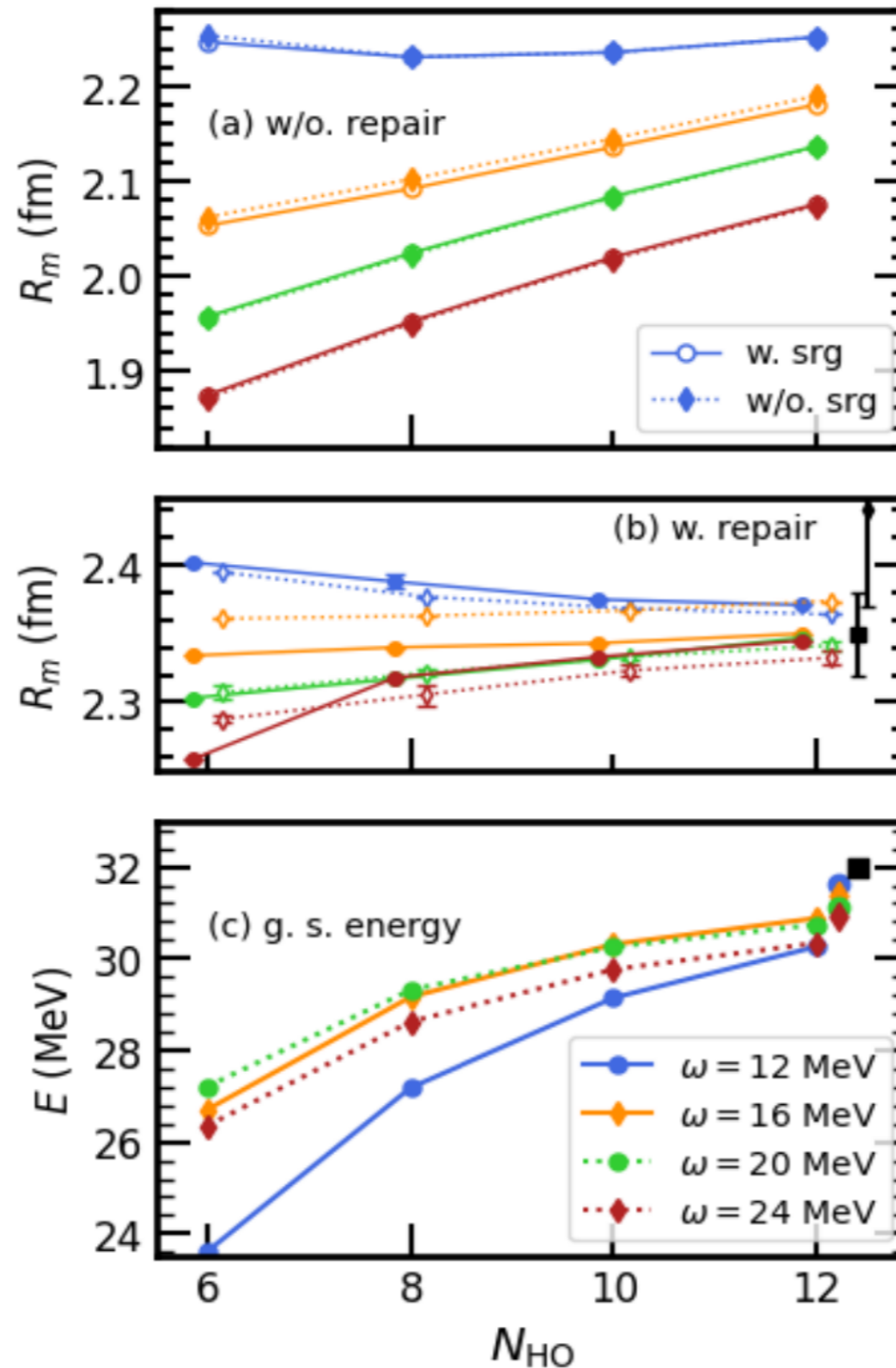
## ${}^6\text{Li}$ : fitting the correction

- fit between 77 ranges between 2.5 and 4.8 fm for different  $N_{HO}$  and selected  $\omega$
- choose densities that give same value at 7 fm for each  $N_{HO}$  within tolerance
- lower tolerance until radii are the same



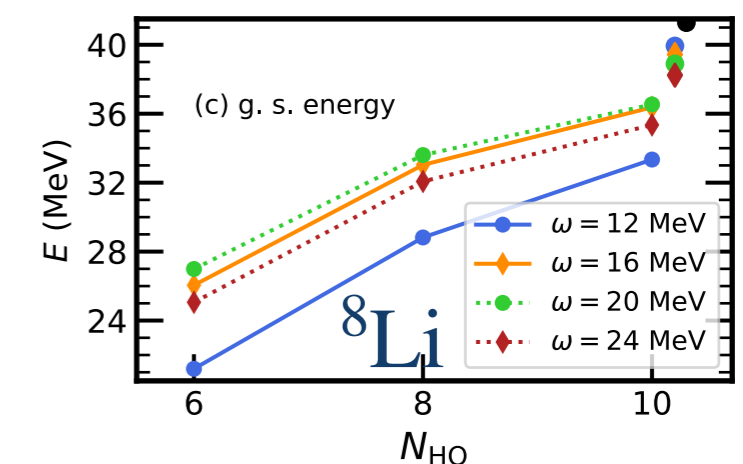
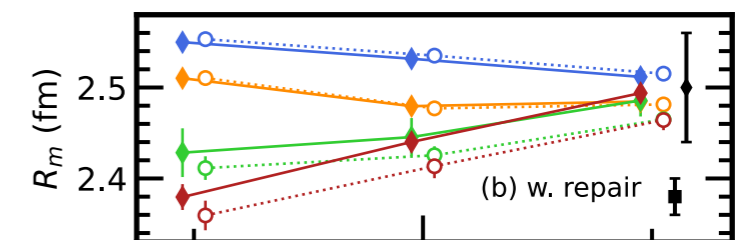
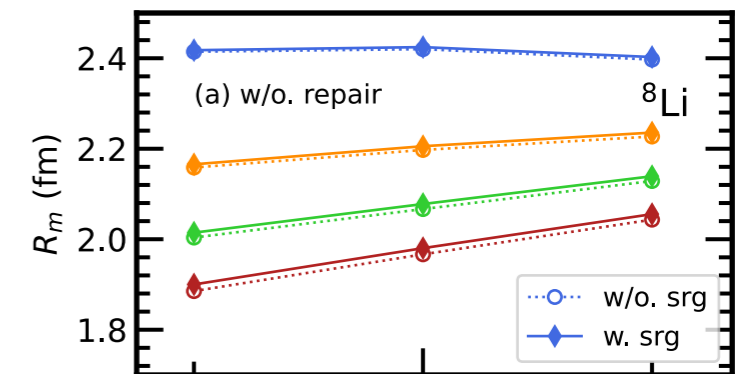
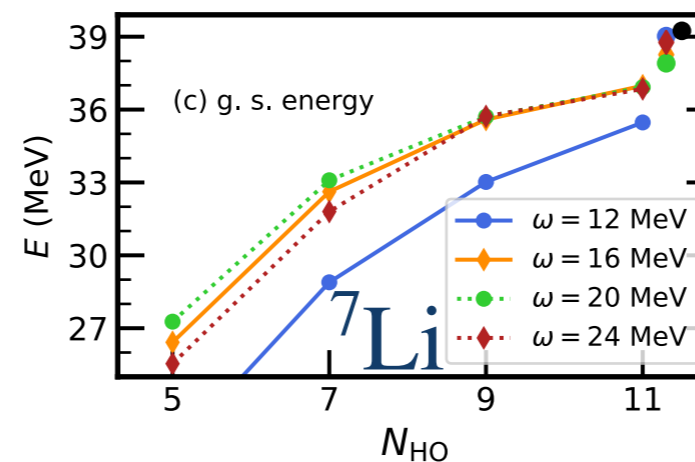
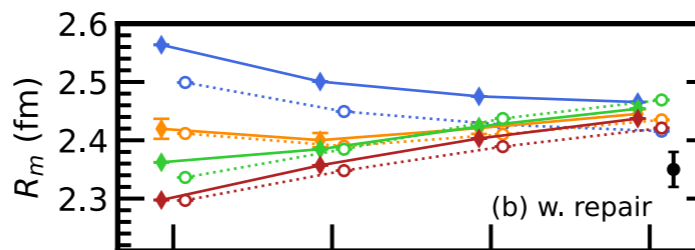
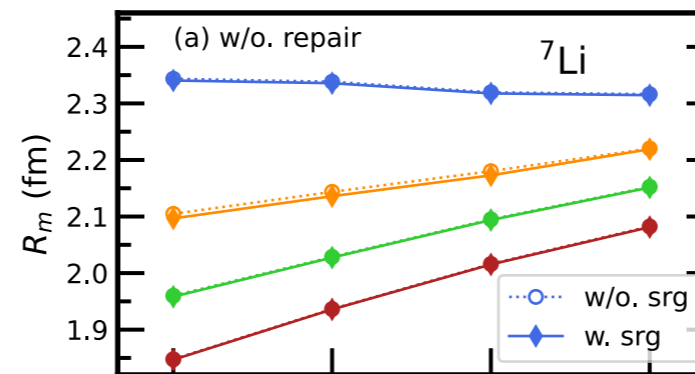
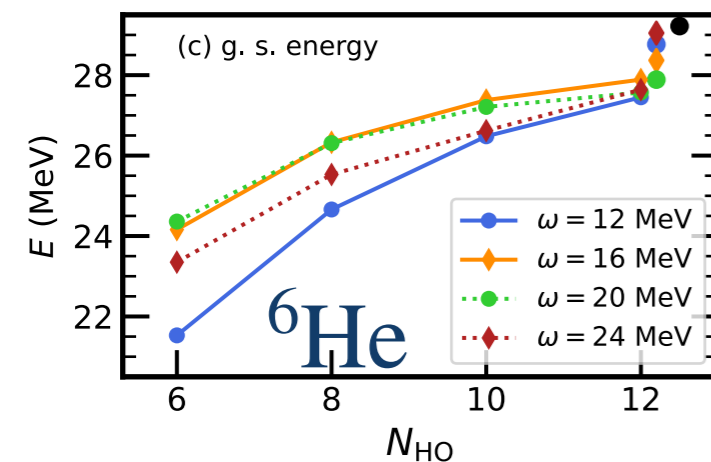
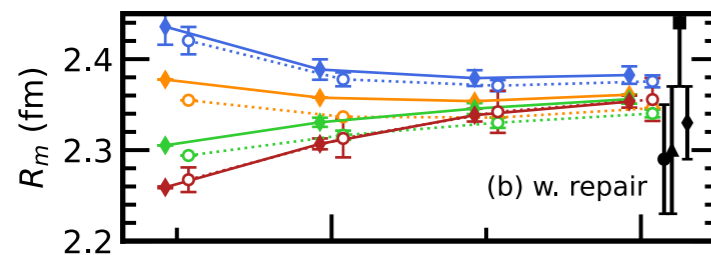
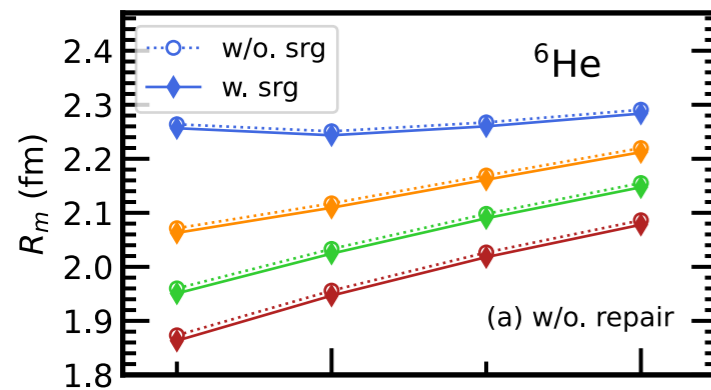
# Two-nucleon densities

- SRG correction for radius small
- tail correction important to obtain convergence
- matter radius consistent with experiment



# Two-nucleon densities

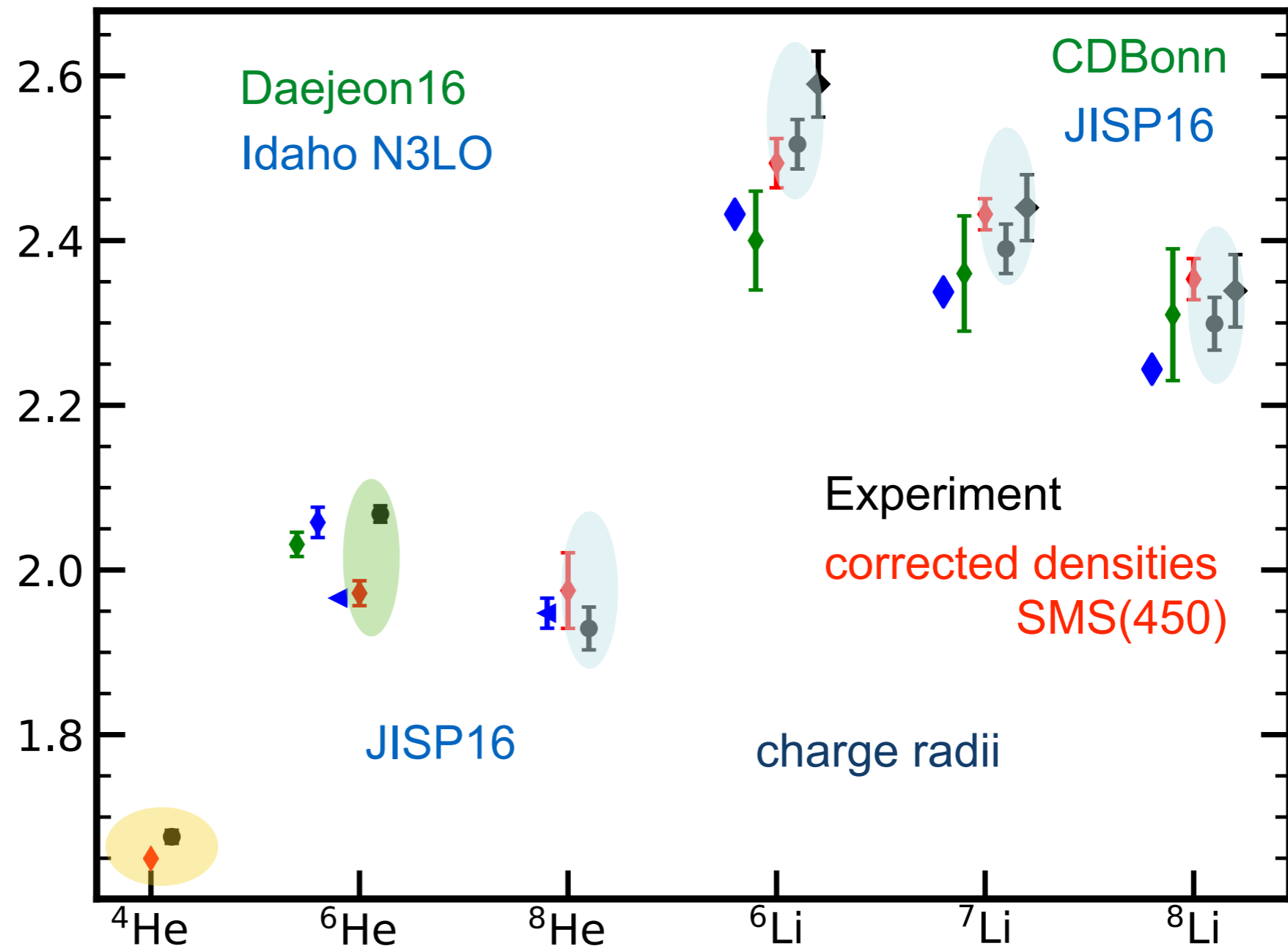
- without correction no convergence and difficult extrapolations
- correction leads to  $\omega$  **independent** result
- **radii increase** due to correction
- generally agreement with experiment with large uncertainties



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# Two-nucleon densities

- does not include 2N corrections (see Filin et al. PRL 2020)
- also **charge radii generally increase** due to correction
- mostly agreement with experiment with large uncertainties



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- **YN interactions not well understood**
  - *scarce YN data*
  - *more information necessary to solve "hyperon puzzle"*
- **Hypernuclei provide important constraints**
  - $^1S_0$   $\Lambda N$  scattering length &  $^3_{\Lambda}\text{H}$
  - $^1S_0$   $\Lambda\Lambda$  scattering length &  $^6_{\Lambda\Lambda}\text{He}$  & predictions for  $A=4,5$
  - CSB of  $\Lambda N$  scattering &  $^4_{\Lambda}\text{He}$  /  $^4_{\Lambda}\text{H}$
- **New SMS YN interactions**
  - *order LO, NLO and N<sup>2</sup>LO allow uncertainty quantification*
  - *have a **non-unique** determination of contact interactions (data necessary)*
- **Chiral 3BF**
  - *decuplet saturation alone does not improve spin dependence*
  - *spin-dependent  $\Lambda\text{NN}$  leads to further improvement*
  - *study cutoff dependence / application to more p-shell hypernuclei*
- **SRG & long-range correction to densities**
  - *increased accuracy of densities*
  - *new applications of NCSM wave functions possible*
  - *form factor calculations in progress (including 2N charge densities)*